

# CO(1-0) EMISSION FROM QUASAR HOST GALAXIES BEYOND REDSHIFT 4

*From Z-Machines to ALMA:  
(Sub) Millimeter Spectroscopy of Galaxies*

January 14, 2006

**Dominik Riechers**

Max-Planck-Institute for Astronomy, Heidelberg, Germany



Deutsche  
Forschungsgemeinschaft

**DFG**

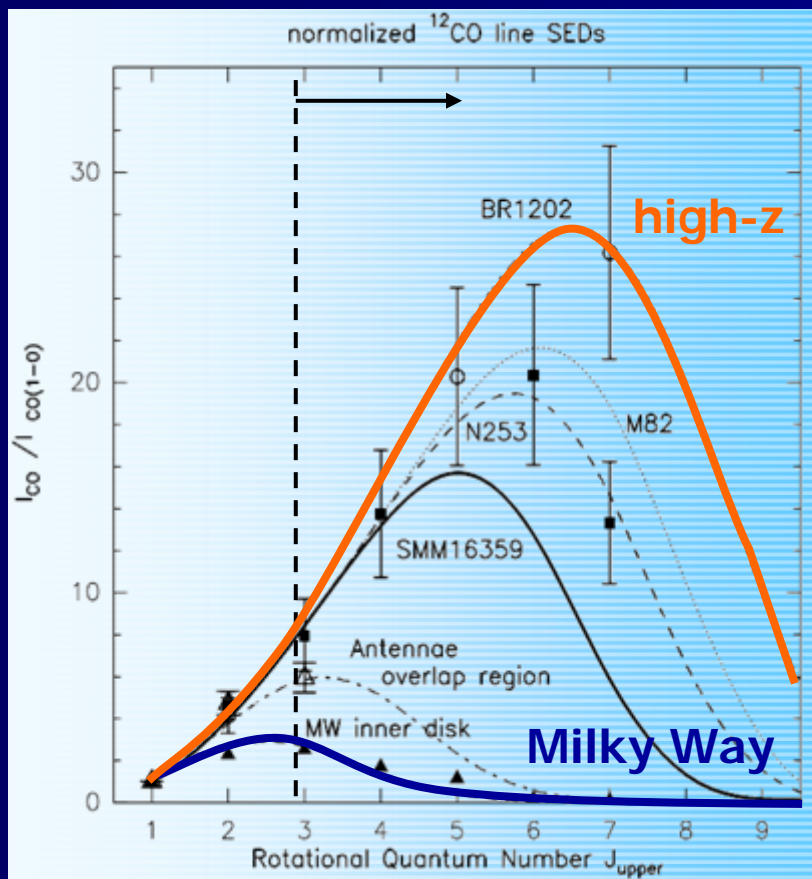
Priority Programme 1177

---

F. Walter (MPIA), C. L. Carilli (NRAO), K. K. Knudsen (MPIA), K. Y. Lo (NRAO),  
D. Benford, J. Staguhn (NASA/GSFC), T. Hunter (CfA), F. Bertoldi (Univ. of Bonn),  
C. Henkel, K. Menten, A. Weiss (MPIfR), M. S. Yun (UMASS), N. Z. Scoville (Caltech)

---

# HIGH-Z CO OBSERVATIONS



- Typically observed at high-z: CO(3-2) and higher

But:

- These obs would miss a low excitation 'cold' component like e.g. found in the Milky Way
- only CO(1-0) traces full  $M(\text{H}_2)$

# 100M SINGLE-DISH TELESCOPES

CO(1-0) observations in high-z QSOs are possible in K band



GBT

**GBT:** 1.6 GHz bandwidth (dual polarization)

- 22000 km/s @22GHz ( $\Delta z/z=0.09$  @z=4.2)
- Future: Zpectrometer, 14 GHz bandwidth @26-40 GHz

22 GHz: CO(1-0) @z=4.2

**Effelsberg:** 500 MHz bandwidth (dual pol.)

- 6000 km/s @22GHz ( $\Delta z/z=0.03$  @z=4.2)

⇒ accurate redshift of QSO host not known from optical/IR, but bandwidth large enough to account for difference to CO redshift

Comparison to **VLA:** 50 MHz bandwidth

- 600 km/s @22GHz ( $\Delta z/z=0.003$  @z=4.2)

⇒ **EVLA** – up to 8 GHz bandwidth per pol.



Effelsberg



# SAMPLE SELECTION: HIGH-Z CO DETECTIONS

■ SMMJ02396-0134	1.062	SMM	■ SMMJ04135+10277	2.846	OSO
■ Q0957+561	1.414	RLQ	■ B3J2330+3927	3.094	RG
■ HR10	1.439	ERO	■ SMMJ22174+0015	3.099	SMM
■ IRAS10214+4724	2.286	OSO	■ MG0751+2716	3.200	RLQ
■ SMMJ16371+4053	2.380	SMM	■ SMMJ09431+4700	3.346	SMM
■ SMMJ16368+4057	2.385	SMM	■ SMMJ13120+4242	3.408	SMM
■ 53W002	2.394	RG	■ TNJ0121+1320	3.520	RG
■ SMMJ16366+4105	2.450	SMM	■ 6C1909+722	3.532	RG
■ SMMJ04431+0210	2.509	SMM	■ 4C60.07	3.791	RG
■ SMMJ16359+6612	2.517	SMM	■ 4C41.17	3.796	RG
■ Cloverleaf	2.558	OSO	■ APM08279+5255	3.911	RQQ
■ SMMJ14011+0252	2.565	SMM	■ PSSJ2322+1944	4.119	OSO
■ VCVJ1409+5628	2.583	OSO	■ BRI1335-0417	4.407	OSO
■ MG0414-0534	2.639	RLQ	■ BRI0952-0115	4.434	OSO
■ MS1512-cB58	2.727	LBQ	■ BR1202-0725	4.693	OSO
■ LBQS1230+1627B	2.741	OSO	■ TNJ0924-2201	5.202	RG
■ RXJ0911+0551	2.796	RQQ	■ SDSSJ1148+5251	6.419	OSO
■ SMMJ02399-0136	2.808	SMM			

PdBI



- List of all detections (1991-2005)
- 1 – 3 mm, mostly CO(3-2), CO(4-3) and higher

see e.g. Solomon & Vanden Bout (2005)



# SAMPLE SELECTION: HIGH-Z CO DETECTIONS

■ -				■ SMMJ04135+10277	2.846	QSO
■ Q0957+561	1.414	RLQ		■ -		
■ -				■ -		
■ IRAS10214+4724	2.286	QSO		■ MG0751+2716	3.200	RLQ
■ -				■ -		
■ -				■ -		
■ -				■ -		
■ -				■ -		
■ -				■ -		
■ Cloverleaf	2.558	QSO		■ APM08279+5255	3.911	RQQ
■ -				■ PSSJ2322+1944	4.119	QSO
■ VCVJ1409+5628	2.583	QSO		■ BRI1335-0417	4.407	QSO
■ MG0414-0534	2.639	RLQ		■ BRI0952-0115	4.434	QSO
■ MS1512-cB58	2.727	LBQ		■ BR1202-0725	4.693	QSO
■ LBQS1230+1627B	2.741	QSO		■ -		
■ RXJ0911+0551	2.796	RQQ		■ SDSSJ1148+5251	6.419	QSO
■ -						

QSOs only!



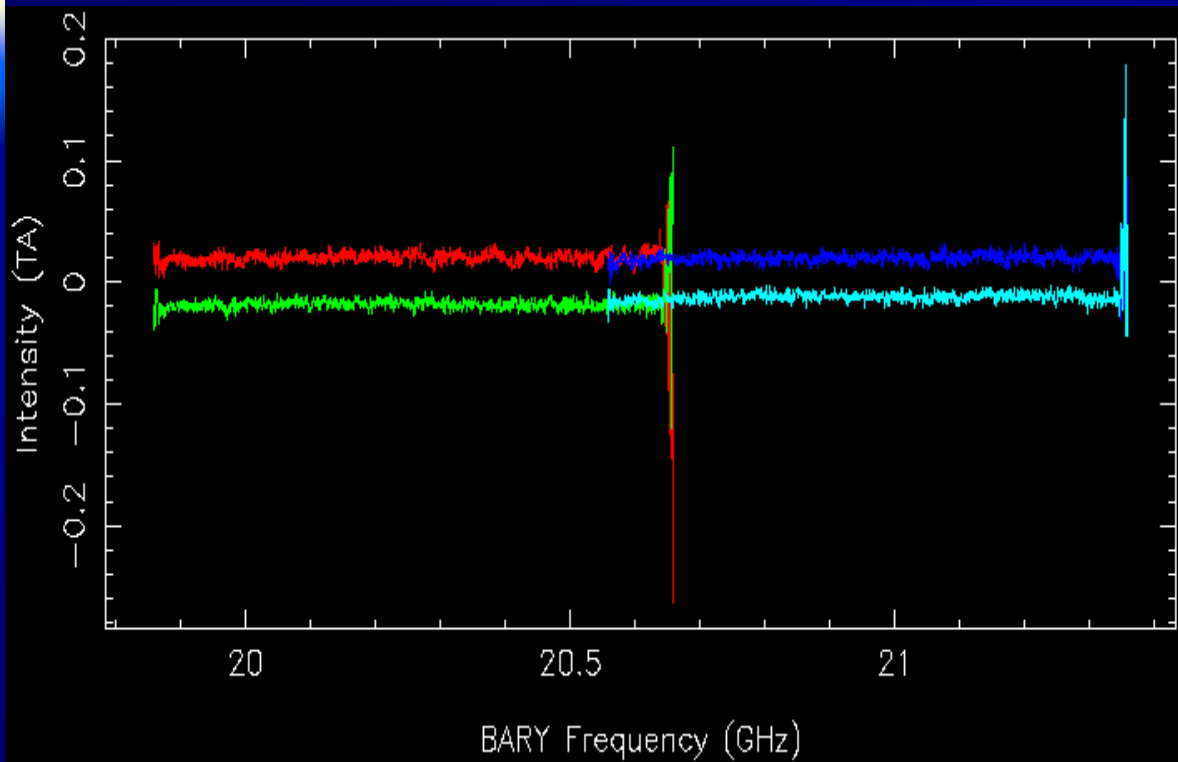
# GBT: DATA QUALITY ASSESSMENT

- 90 hours, 3 QSOs in K band
- Position Switched Dual Beam observations
- Mostly excellent winter weather in December 2004:

$$T_{\text{sys}} = 20\text{-}25 \text{ K}$$



# GBT: DATA QUALITY ASSESSMENT

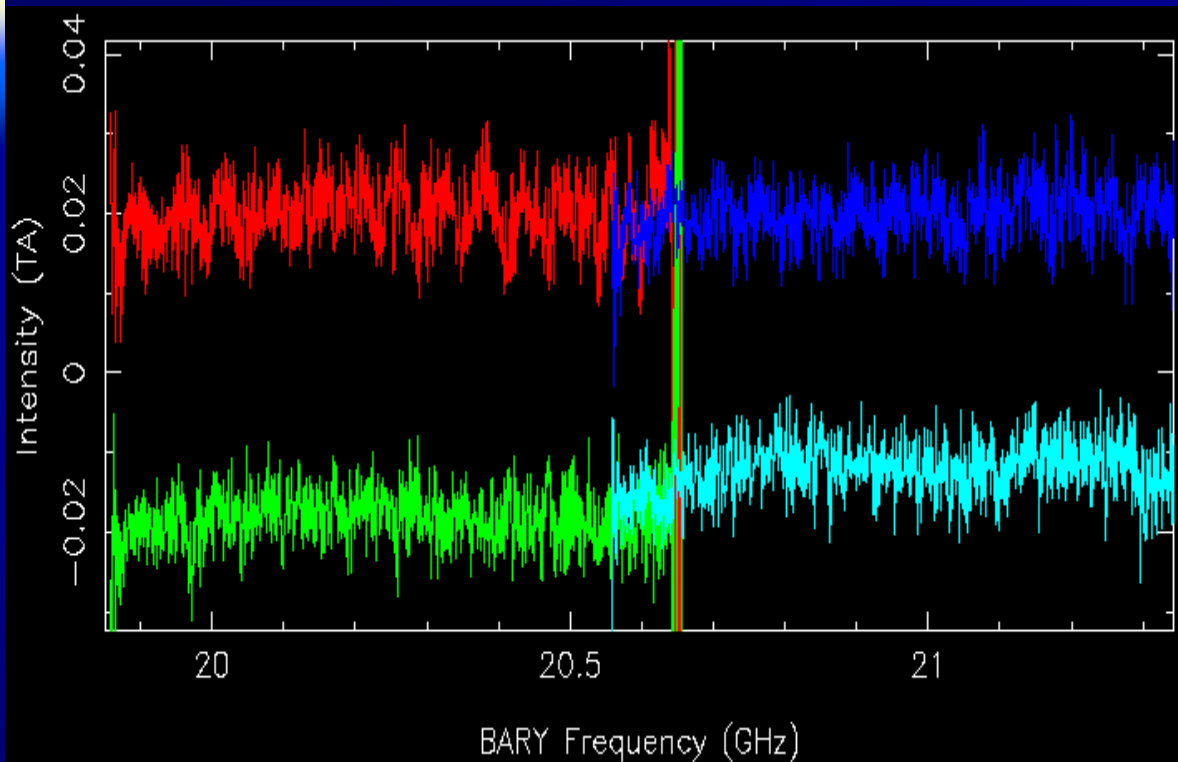


Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 22 \text{ K}$ , 1.6 GHz, LL/RR polarization



# GBT: DATA QUALITY ASSESSMENT

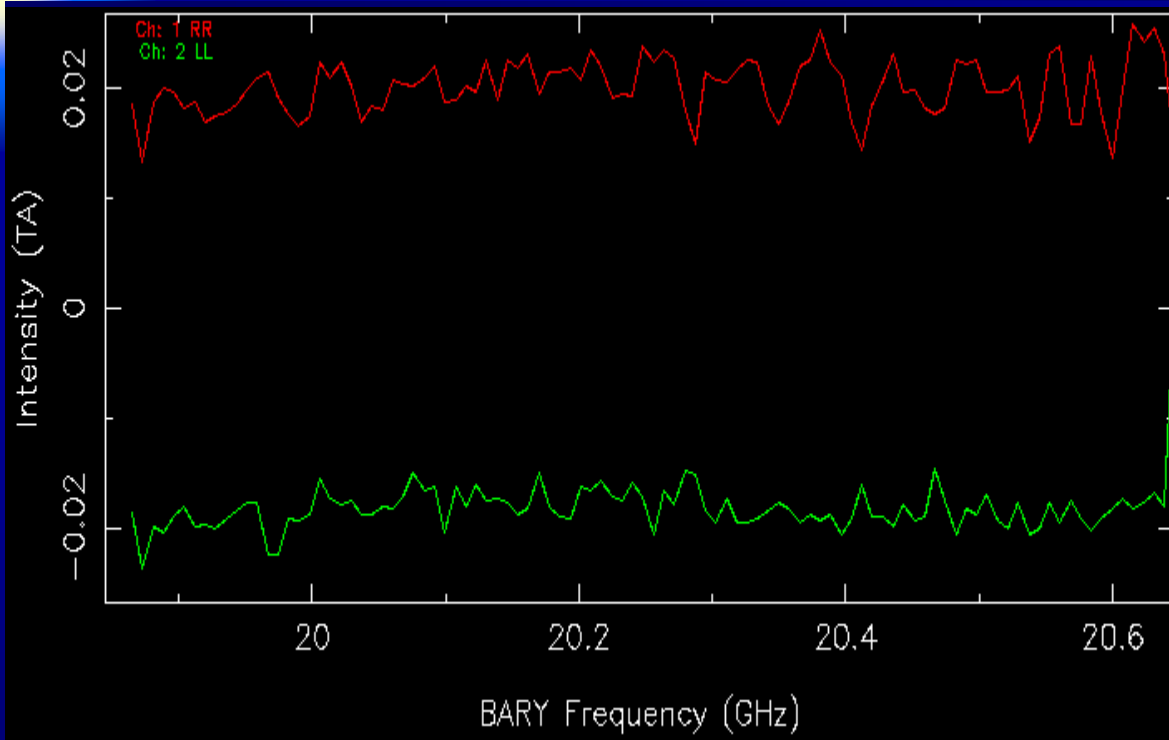


Zoomed in

Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 22 \text{ K}$ , 1.6 GHz, LL/RR polarization

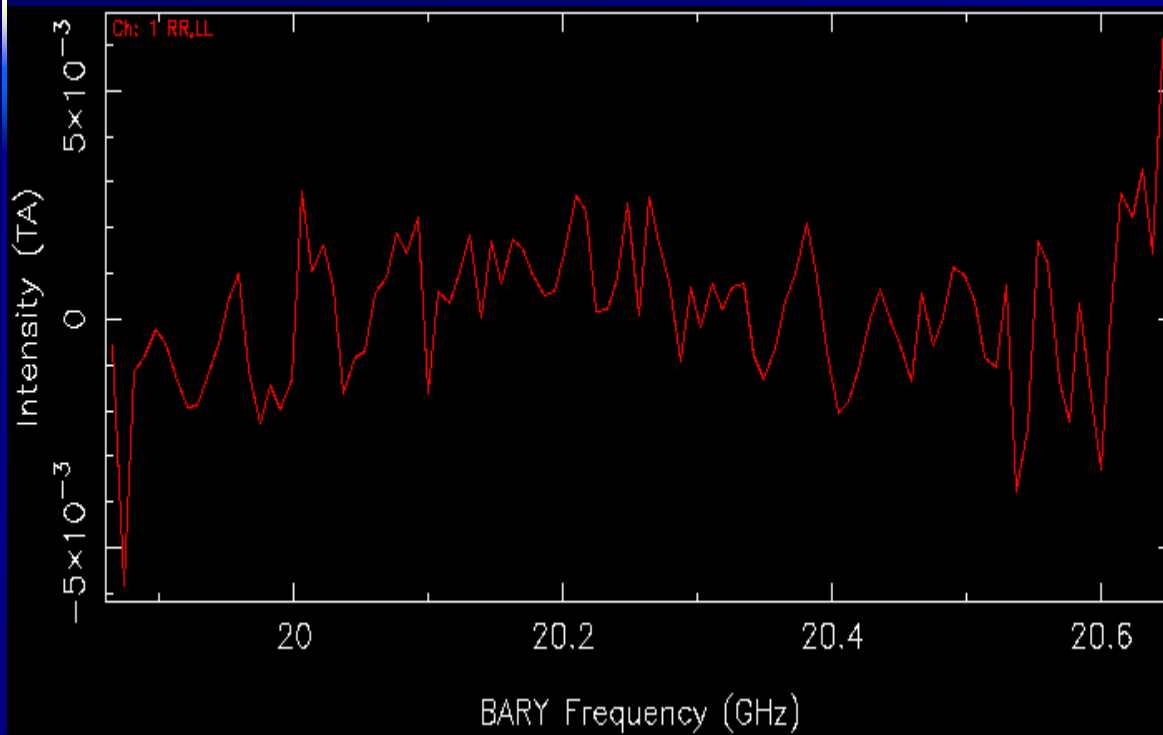
# GBT: DATA QUALITY ASSESSMENT



↑ Zoomed in, IF1,  
Boxcar 20 ch  
Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 22 \text{ K}$ , 0.8 GHz, LL/RR polarization

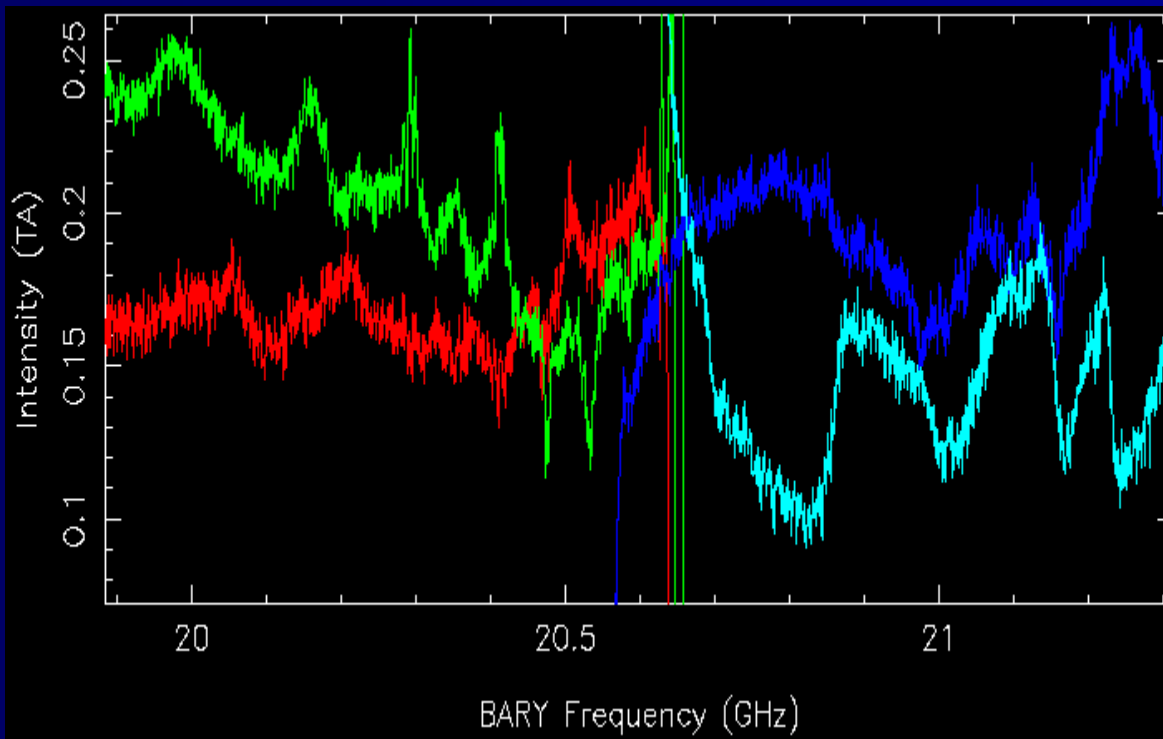
# GBT: DATA QUALITY ASSESSMENT



↑ Zoomed in, IF1,  
Box20, avg. pol.  
Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 22$  K, 0.8 GHz, (LL+RR) polarization

# GBT DATA: THE DARK SIDE

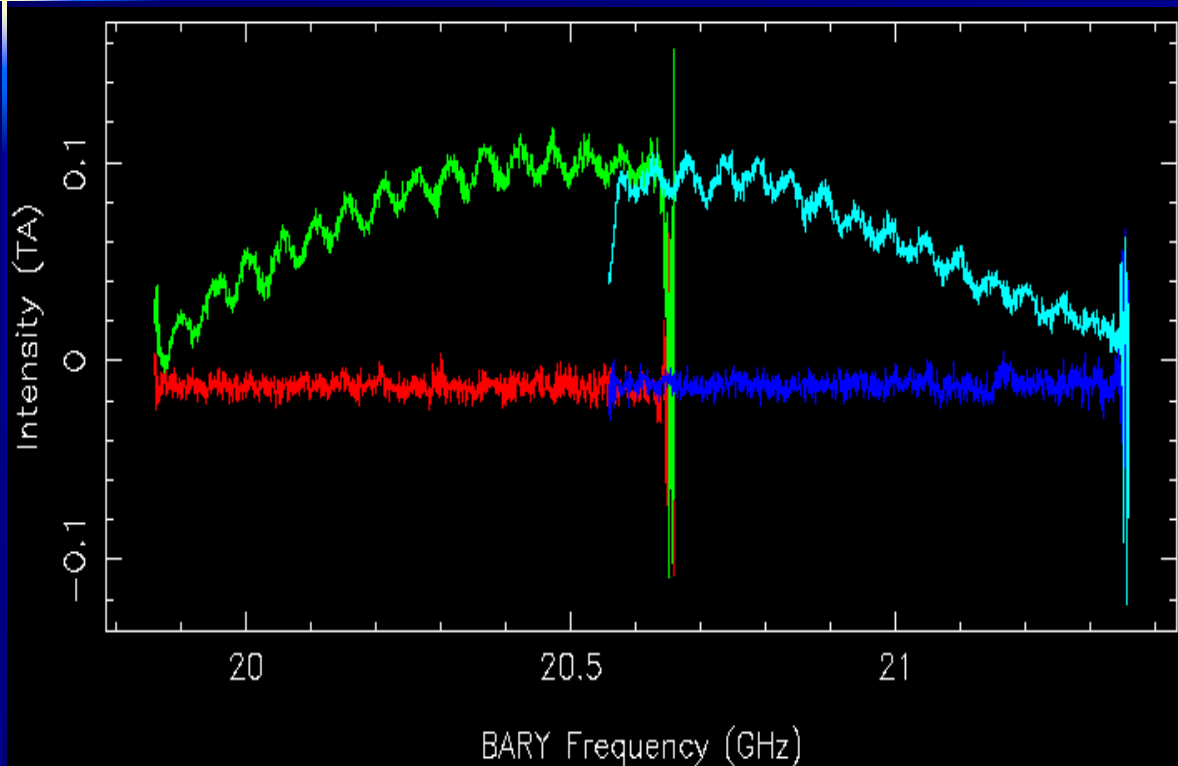


Weather

Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 40 \text{ K}$ , 1.6 GHz, LL/RR polarization

# GBT DATA: THE DARK SIDE

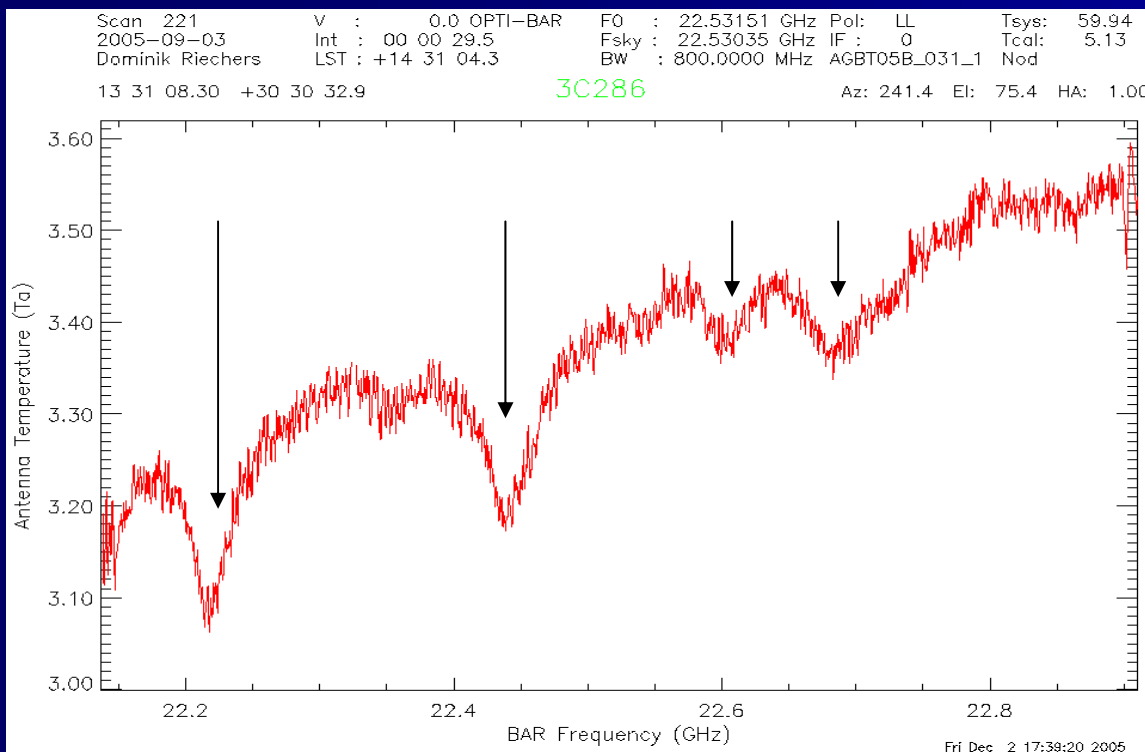


Resonance  
Features

Gain: 1.5K/Jy

90 s integration,  $T_{\text{sys}} = 40$  K, 1.6 GHz, LL/RR polarization

# GBT DATA: THE DARK SIDE

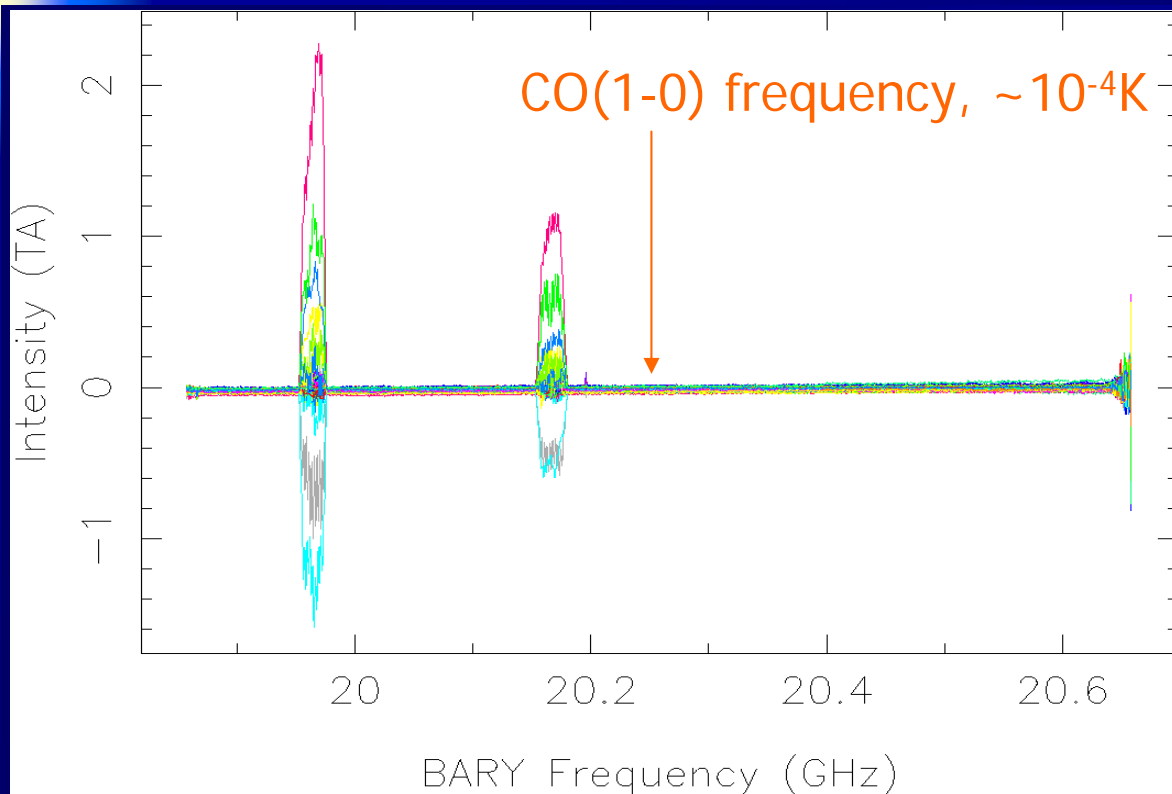


Birdies

Gain: 1.5K/Jy

30 s integration,  $T_{\text{sys}} = 60$  K, 0.8 GHz, LL/RR polarization

# GBT DATA: THE DARK SIDE



Military Satellite

Gain: 1.5K/Jy

20x90 s integration, 0.8 GHz, LL/RR polarization

# GBT DATA: THE DARK SIDE

- Fixable: Satellite (spectral baselines remain stable)
- Flaggable: Weather, Resonances
- 'Deadly': Birdies





# GBT: DATA REDUCTION

## aips++

- standard calib (late 2004)
- standard calib (early 2005)
- Bob Garwood's calib
- Ron Maddalena's calib

## IDL

- Glen Langston's routines
- GBTIDL

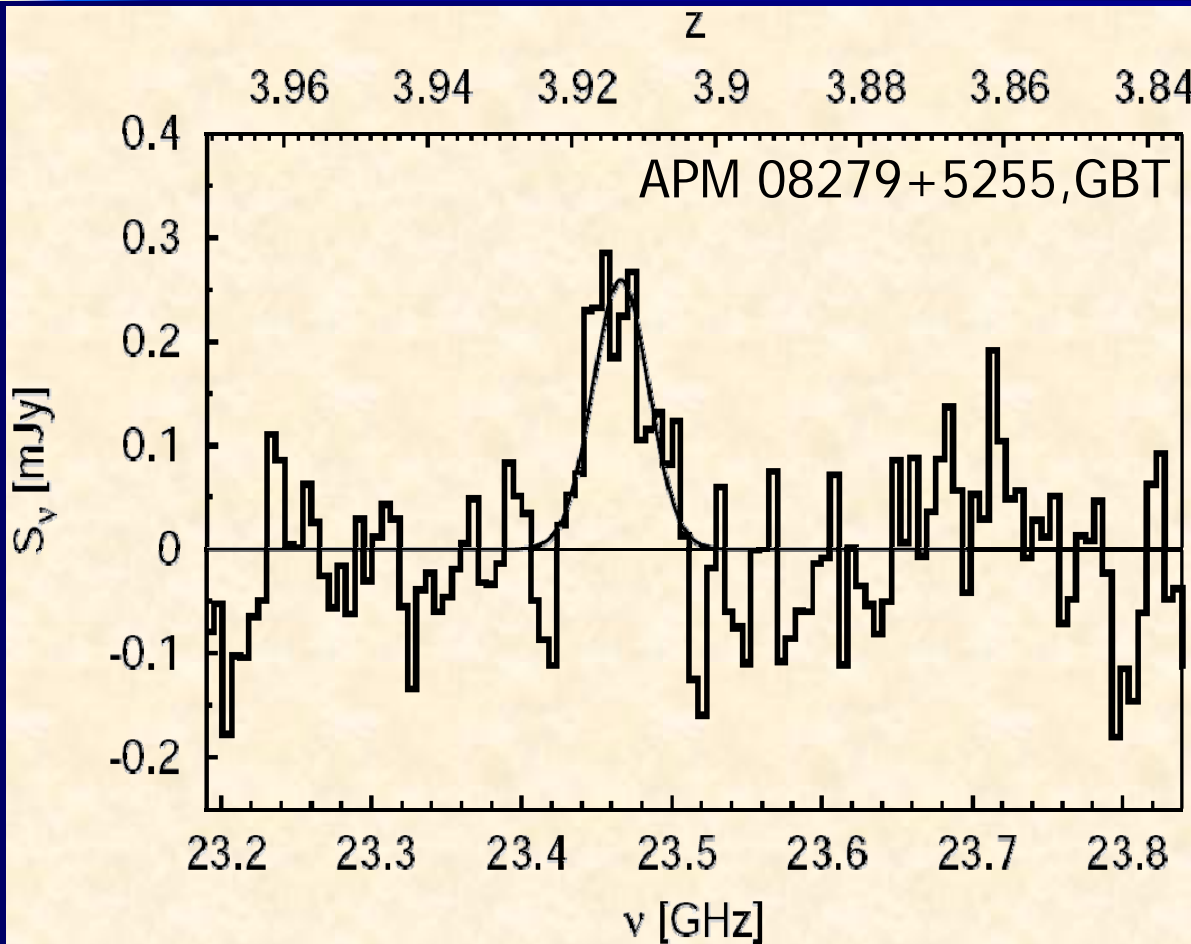
- Vanden Bout, Solomon & Maddalena (2004) scheme

[HCN(1-0) in IRAS F10214+4724 ( $z=2.3$ )]

**DETECTION OF CO(1-0) @ $z>4$ !**

# CO(1-0): SPECTRA

APM 08279+5255 ( $z=3.91$ )



## GBT

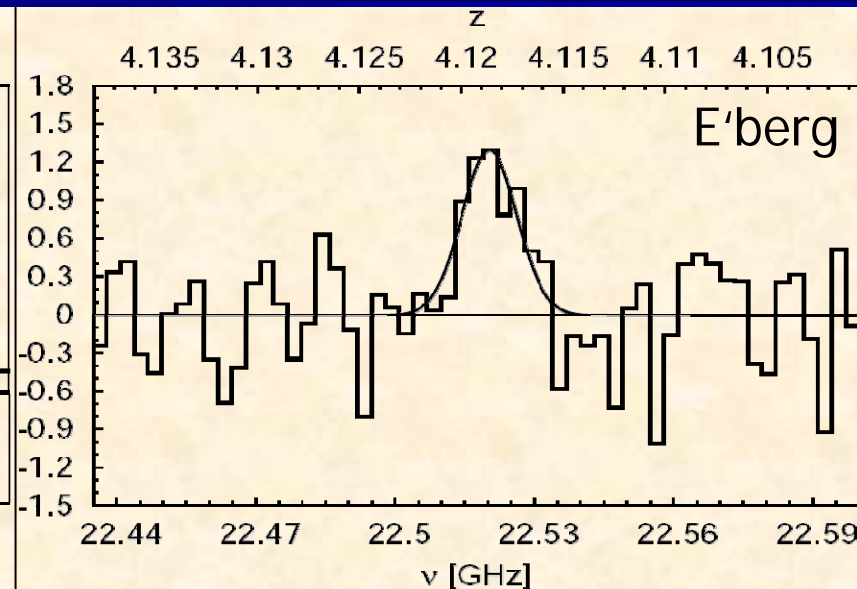
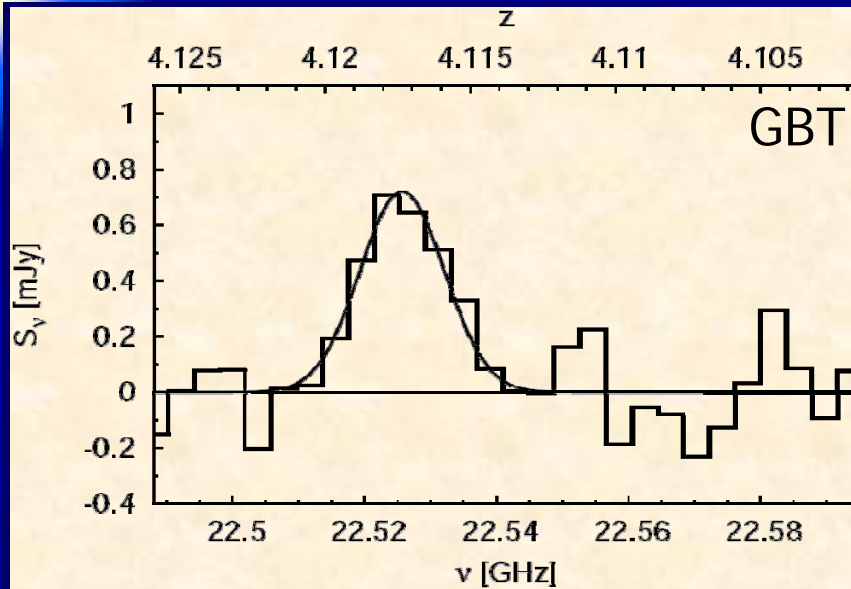
- 35 hours total
- 22 hours on src
- rms: 80  $\mu$ Jy
- 75 km/s res ( $\sigma$ )

- 280  $\mu$ Jy peak
- 560 km/s FWHM

2nd order polynomial fit to spectral baseline subtracted

# CO(1-0): SPECTRA

PSS J2322+1944 ( $z=4.12$ )



## GBT

- 23 hours total
- 15 hours on src
- rms: 160  $\mu$ Jy
- 52 km/s res ( $\sigma$ )

## Effelsberg

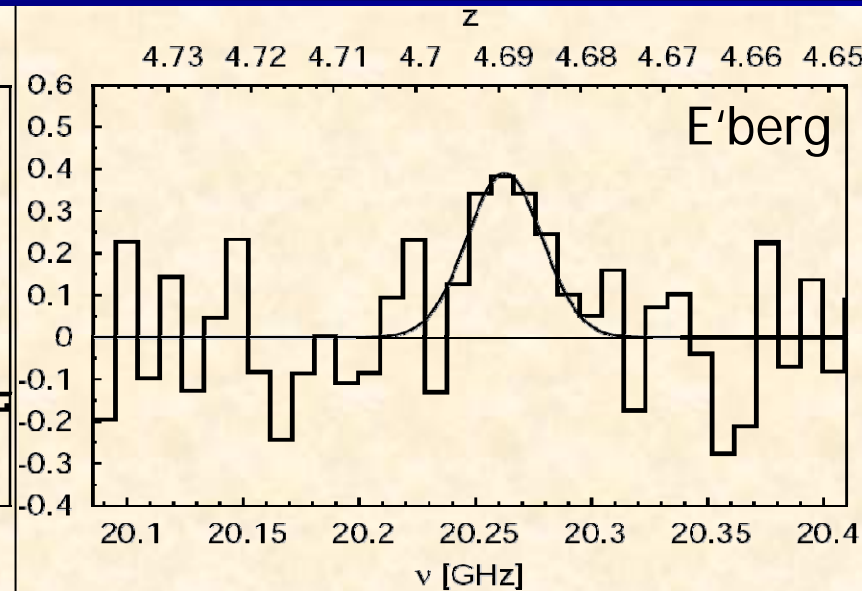
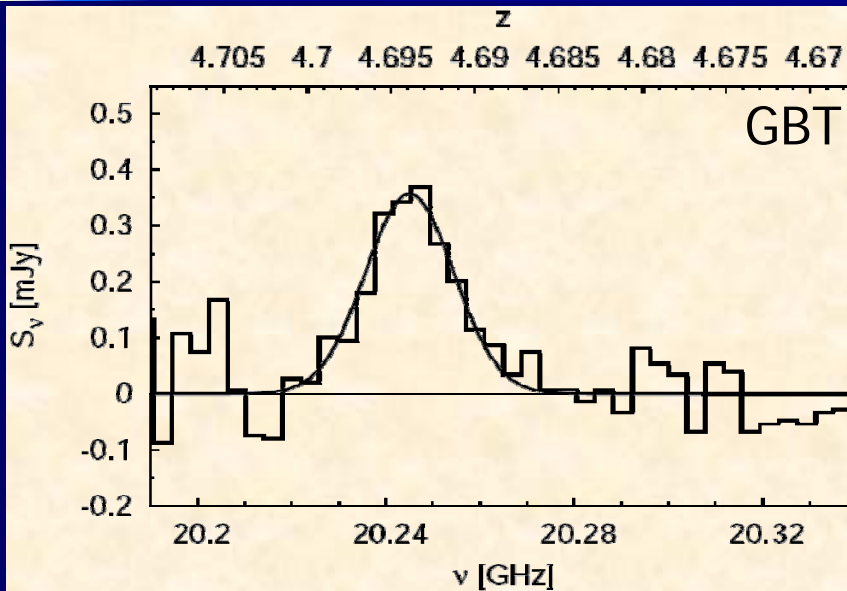
- 20 hours total
- 8 hours on src
- rms: 460  $\mu$ Jy
- 40 km/s res ( $\sigma$ )

- 720  $\mu$ Jy peak
- 202 km/s FWHM

linear fits to spectral baseline subtracted

# CO(1-0): SPECTRA

BR 1202-0725 ( $z=4.69$ )



## GBT

- 32 hours total
- 20 hours on src
- rms: 70  $\mu$ Jy
- 58 km/s res ( $\sigma_{sm}$ )

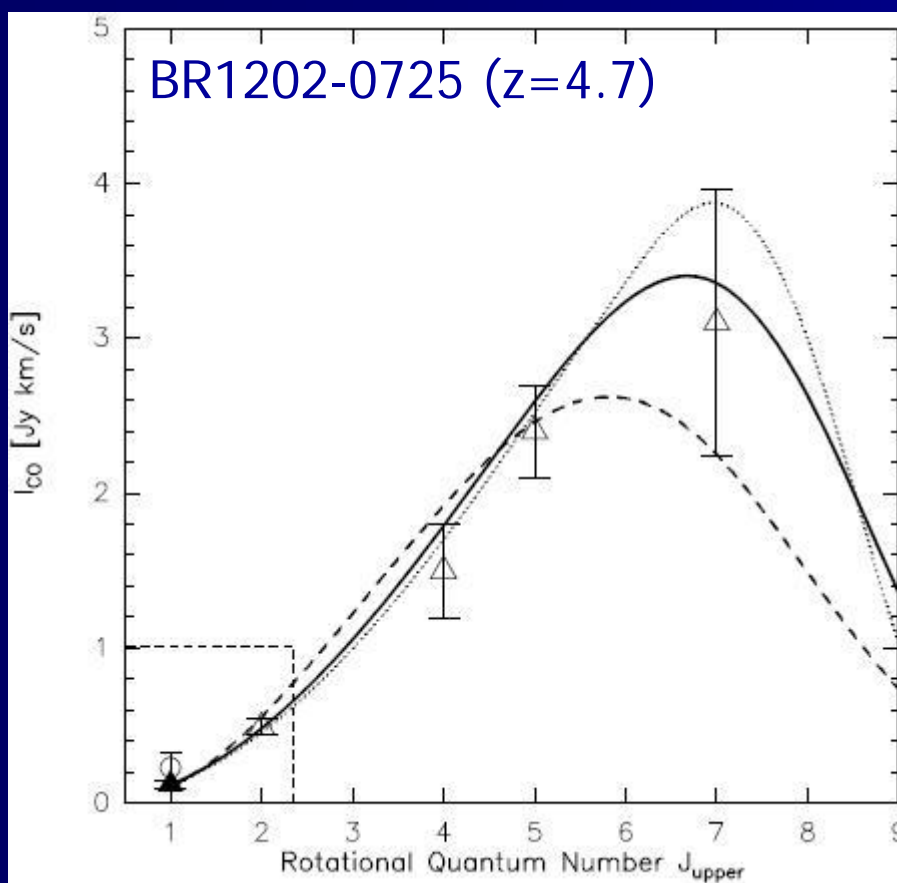
## Effelsberg

- 20 hours total
- 8 hours on src
- rms: 160  $\mu$ Jy
- 141 km/s res ( $\sigma_{sm}$ )

- 360  $\mu$ Jy peak
- 329 km/s FWHM

order 2/1 fit to spec. baseline subtracted

# EXCITATION CONDITIONS



Large Velocity Gradient (LVG) modeling (1 component) on higher- $J$  transitions

- $T_{kin} = 60$  K,
- $n(H_2) = 10^{4.1} \text{cm}^{-3}$
- all CO(1-0) flux recovered

⇒ almost fully thermalized

( $I_{CO} \sim v^2$ ) up to CO(4-3)

⇒ traces as much gas as CO(1-0)

# SUMMARY

- First-time detection of CO @ $z>4$  with 100m single-dish telescopes
- caveat: spectral baselines still limit the detectability
  - ⇒ detection of CO in new, possibly fainter sources very difficult
  - ⇒ requirement for success of observations: the very best weather

- 
- massive reservoirs of molecular gas:  
( $M(\text{H}_2) \sim 10^{10} M_\odot$ )
  - no evidence for additional massive, cold CO(1-0) components which are not visible at high- $J$  CO transitions

