POLARIZATION SCIENCE WITH CARMA

PROBING SMALL-SCALE MAGNETIC FIELDS IN STAR-FORMING REGIONS

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Special Session: Polarization ALMA 2013 Rocks! Kona, Hawai'i

Outline

• CARMA

• 1mm polarization system

Science results & future proposals

- TADPOL survey: protostellar polarization
- Circumstellar disk polarization & formation
- Other projects: Serpens, SgrA*

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CARMA

Combined Array for Research in Millimeter-wave Astronomy





- 6 ×10-m, 9 × 6-m, 8 × 3.5-m telescopes

 Observations at 1 cm, 3 mm, and 1 mm (polarization!)

 Located in Cedar Flat, CA (near Bishop)

1 mm dual-polarization receivers





Plambeck & Engargiola, CARMA Memo #54

Polarizer simulation

\leftarrow Feed horn (sky)

$\text{OMT} \not \rightarrow$

Plambeck & Engargiola, CARMA Memo #54



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Misalignment of B-fields and outflows





TADPOL collaboration

• UC Berkeley

Chat Hull (PI), Dick Plambeck, Mel Wright, Carl Heiles, Geoff Bower

• University of Maryland

Marc Pound, Alberto Bolatto, Katherine Jameson, Lee Mundy

• Caltech

Thushara Pillai, John Carpenter, James Lamb, Nikolaus Volgenau

• University of Illinois, Urbana-Champagne

Ian Stephens, Leslie Looney, Woojin Kwon, Dick Crutcher, Nick Hakobian

• Other

Dan Marrone (Arizona), Meredith Hughes (Wesleyan), John Vaillancourt & Göran Sandell (USRA-SOFIA), John Tobin (NRAO), Jason Fiege (Manitoba), Erica Franzmann (Manitoba), Martin Houde (UWO, Caltech), Brenda Matthews (NRC-CNRC)





35 sources

Triples number of interferometric polarization maps

~300 observing hours CARMA C, D, & E arrays

1 - 4'' resolution

10× higher resolution than CSO & JCMT Probes intermediate region between ~0.1 pc (single-dish) and ~100 AU (ALMA)



Credit: Bill Saxton, Harvard Smithsonian Center for Astrophysics



RA offset (arcsec; J2000)





NGC 1333-IRAS 2A



TADPOL results Ser-emb 8 1510DEC offset (arcsec; J2000) ß 0 Ω -10 15 5 0 $^{-5}$ -10 10

RA offset (arcsec; J2000)



Upper limits on circumstellar disk polarization



• CARMA + SMA results

 Disks are not strongly polarized at ~100 AU scales:

P < 0.5%

Hughes, Hull, Wilner, & Plambeck 2013, AJ, 145, 115

Implications for disk formation



- Weak & misaligned Bfields in cores could aid disk formation
- Addresses "magnetic braking catastrophe"
- Points are from simulations in Joos+ 2012
- We find that ~10–50% of Class 0/I stars should form Keplerian disks

Krumholz, Crutcher, & Hull 2013, ApJL, 767, L11

Other projects

• Filamentary B-field structure in Serpens

• Observing same field as ALMA (Mundy *et al.*)

Galactic center (SgrA*)

- Rotation measure, simultaneous with SMA
- G2 cloud

Summary

- CARMA 1 mm polarization system is fully functional, and accepting proposals (next deadline is **next month** (May, '13))
- Wide array of science
- TADPOL results: B-fields are either **preferentially misaligned** (perp.) or **randomly aligned** with respect to outflows at the ~1000 AU scale
 - Thus, circumstellar disks are misaligned with fields in the cores from which they formed
- TADPOL results: <u>arXiv:1212.0540</u>
- TADPOL survey (CARMA key project): <u>tadpol.astro.illinois.edu</u>
- Questions? Email: <u>chat@astro.berkeley.edu</u>

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