

**Summary of Results**  
**ALMA Front End Cryogenics and Optics Meeting**  
**26 and 27 September 2001**  
**Rutherford Appleton Laboratory**

09 October 2001, W. Wild

## **1. General**

On 26 and 27 September, a series of ALMA front end technical meetings was held at Rutherford Appleton Laboratory. These included

- Cryogenics and optics meeting (organized by M. Carter and A. Orłowska)
- Mixer meeting (organized by B. Lazareff)
- Photonics LO meeting (organized by B. Ellison)

At the same time, small scale WVR and antenna meetings were held, and on 28 and 29 September, an E-AEC meeting was held.

This note summarizes the results of the cryogenics and optics meeting.

## **2. Agenda and Attendees**

The agenda and list of attendees are given in Annex B and C, respectively.

## **3. Purpose of the meeting**

The aims and objectives of the cryogenics and optics meeting were

- to present the current status of the cryostat and optics design
- to give RAL the "go ahead" for the construction of a prototype cryostat (engineering model), in particular
  - reach agreement on overall concept
  - confirm design solutions (avoid later re-design)
  - identify problems and outstanding technical issues
  - Discuss and try to resolve outstanding issues
- to become aware of anything in the FE design affecting cryogenics and optics

## **4. Main results**

The meeting was held in a very collaborative atmosphere, and open and fruitful discussions took place. Basically all objectives have been reached during the 1.5 day meeting. Where issues could not be resolved during the meeting, deadlines for corresponding action items have been set (see below). The main results are:

- The RAL cryostat design is advanced enough to start building a prototype (engineering model).

- An electromagnetic analysis of the optics design has been carried out by MRAO, and a detailed report was presented (available at: <http://www.mrao.cam.ac.uk/~cytham/>).
- It was decided at the meeting to have identical mechanical cartridge interfaces to the RAL prototype cryostat and the Japanese ASTE cryostat (although this means rebuilding parts of the ASTE cryostat, and some redesign of the RAL cryostat). In principle, this makes changing cartridges between the RAL and ASTE cryostats possible.
- The RAL prototype cryostat and ASTE cryostat will have different thermal links. This was not considered to be a problem as long as the cartridges are interchangeable.
- Three different cartridge structures were proposed, the so-called "cylinder" and "truss" structures by RAL, and a "column" structure by NAOJ. Their pro and cons were discussed to some extent, but it was felt that no decision between the structures is needed now. Fabrication, use and measurements with different structures would give necessary data for taking a decision.
- The following clear windows are foreseen: plastic lenses (bands 1 and 2), quartz (bands 3 to 6), TBD – possibly quartz (bands 7 to 10). Calculations for the high freq bands are pending.
- The cryostat thermal model includes a number of wires (see Annex A), and the cartridge groups are asked to review the assumptions and send comments to RAL.
- The so-called "70 K stage" can have a temperature of 85 +/- 15 K (i.e. between 70 and 100 K).
- The issue of optical alignment and alignment stability was discussed. Further work will be done by B. Lazareff in collaboration with the cartridge and cryostat groups.

## 5. Action items

The following action items (ordered from early to late deadline) were agreed at the meeting. When sending information, please send always a copy to Payne, Sekimoto, and Wild. [Note added in final version: action items A1, A3, A4, and A5 completed.]

Item	By when	Who	Action
A1	08 Oct 2001	M. Harman	<ul style="list-style-type: none"> <li>• Provide new cartridge drawings to cartridge groups</li> </ul>
A2	15 Oct 2001	B. Ellison	<ul style="list-style-type: none"> <li>• collect and distribute electronically meeting viewgraphs to attendees</li> </ul>
A3	15 Oct 2001	G. Ediss	<ul style="list-style-type: none"> <li>• Check with LO multiplier group if new "70 K stage" temperature of 85 K +/- 15 k ok</li> <li>• Send result to A. Orłowska</li> </ul>
A4	15 Oct 2001	G. Ediss	<ul style="list-style-type: none"> <li>• Check with LO multiplier group how much space is needed at 90 K stage for LO components</li> <li>• Results to RAL and cartridge groups <sup>(a)</sup></li> </ul>
A5	15 Oct 2001	G. Ediss	<ul style="list-style-type: none"> <li>• Calculate quartz windows for high frequency bands 7 to 10</li> <li>• Send results to M. Carter</li> </ul>
A6	15 Oct 2001	Cartridge groups <sup>(a)</sup>	<ul style="list-style-type: none"> <li>• Determine minimum distance between 4 K stage and dewar top</li> <li>• Check if 188 mm is ok (ASTE distance)</li> <li>• Send results to M. Carter and A. Orłowska</li> </ul>
A7	15 Oct 2001	B. Lazareff	<ul style="list-style-type: none"> <li>• Release version 1.98 of optics alignment and tolerance report</li> </ul>
A8	31 Oct 2001	Cartridge groups <sup>(a)</sup>	<ul style="list-style-type: none"> <li>• Provide info about needed access to cartridge (size of cut-away holes etc.) to M. Harman</li> </ul>
A9	31 Oct 2001	S. Navarro	<ul style="list-style-type: none"> <li>• Send information about IR filter properties (as far as known) to A. Orłowska</li> </ul>

A10	31 Oct 2001	Cartridge groups <sup>(a)</sup>	<ul style="list-style-type: none"> <li>• Revise A. Orłowska's thermal load sheet (see Annex A) and give feedback to her</li> </ul>
A11	31 Oct 2001	Cartridge groups <sup>(a)</sup>	<ul style="list-style-type: none"> <li>• Inform M. Carter about exact position of dewar window for respective band</li> </ul>
A12	01 Nov 2001	R. Brown	<ul style="list-style-type: none"> <li>• Provide number of windows simultaneously covered with solar filters to M. Carter</li> </ul>
A13	15 Nov 2001	R. Hills, M. Carter	<ul style="list-style-type: none"> <li>• R. Hills to provide WVR size to M. Carter</li> <li>• M. Carter and R. Hills to agree on WVR space within FE widget space</li> </ul>
A14	30 Nov 2001	A. Orłowska	<ul style="list-style-type: none"> <li>• Measure dewar cool down time</li> <li>• Add heat switches and pre-cooling loop(s) to present design</li> <li>• Assess what is required to reach 24 h cool down time</li> </ul>
A15	30 Nov 2001	M. Harman	<ul style="list-style-type: none"> <li>• Explore tube cartridge structure with larger access holes</li> <li>• Perform FE analysis</li> <li>• Results to cartridge groups</li> </ul>
A16	01 Apr 2002	J. Payne	<ul style="list-style-type: none"> <li>• produce band 6 OMT (demonstrate feasibility)</li> </ul>
A17	01 Apr 2002	IRAM	<ul style="list-style-type: none"> <li>• Provide prototype windows to RAL</li> </ul>

<sup>(a)</sup> The following persons should be contacted regarding frequency cartridge issues (the "cartridge groups"):

J. Payne  
G. Ediss  
C. Cunningham  
A. Baryshev  
S. Claude  
V. Belitsky  
Y. Sekimoto  
W. Wild

## Annex A: Included heat loads in RAL thermal model.

(Sheet distributed by A. Orłowska at the cryogenics and optics meeting)

### Thermal model

The thermal model has included the following wiring for each of bands 3-10. We have also assumed 36 mW additional load at 4K for each band that is on. Please let me know if there are major problems with these assumptions.

Anna Orłowska

<b>Interface 12 - 4K</b>	<b>No. of Cables</b>	<b>Type</b>	<b>Heat Load (ESATAN)</b>
SIS Mixer Bias	16 (4 per mixer)	125 $\mu$ m dia. BeCu	0.078mW
Magnet Bias	8 (2 per mixer)	125 $\mu$ m dia. Cu	6.588mW
LNA Bias	20	125 $\mu$ m dia. BeCu	0.097mW
IF Cable	4 (1 per mixer)	0.085 inch SS-SS	0.554mW
W/G LO Input	2	TBD	TBD
4K Temp. Sensor	16 (4 sensors)	125 $\mu$ m dia. BeCu	0.078mW
Stage heater	4 (paired)	125 $\mu$ m dia. BeCu	00.019mW

ALMA SIS Receiver Cartridge Cabling Heat Load Estimate on 4 K Stage

<b>Interface 70 - 12K</b>	<b>No. of Cables</b>	<b>Type</b>	<b>Heat Load (ESATAN)</b>
SIS Mixer Bias	16 (4 per mixer)	125 $\mu$ m dia. BeCu	1.561mW
Magnet Bias	8 (2 per mixer)	125 $\mu$ m dia. Cu	31.003mW
LNA Bias	20	125 $\mu$ m dia. BeCu	1.951mW
IF Cable	4 (1 per mixer)	0.085 inch SS-SS	22.874W
W/G LO Input	2	TBD	TBD
12K Temp. Sensor	8 (2 sensors)	125 $\mu$ m dia. BeCu	0.780mW
Stage heater	4 (paired)	125 $\mu$ m dia. BeCu	0.390mW

ALMA SIS Receiver Cartridge Cabling Heat Load Estimate on 12 K Stage

<b>Interface 300 - 70K</b>	<b>No. of Cables</b>	<b>Type</b>	<b>Heat Load</b>
SIS Mixer Bias	16 (4 per mixer)	125 $\mu$ m dia. BeCu	14.360mW
Magnet Bias	8 (2 per mixer)	125 $\mu$ m dia. Cu	44.940mW
LNA Bias	20	125 $\mu$ m dia. BeCu	17.950mW
IF Cable	4 (1 per mixer)	0.085 inch SS-SS	188.490W
W/G LO Input	2	TBD	TBD
70K Temp. Sensor	8 (2 sensors)	125 $\mu$ m dia. BeCu	7.180mW
Stage heater	4 (paired)	125 $\mu$ m dia. BeCu	3.590mW

ALMA SIS Receiver Cartridge Cabling Heat Load Estimate on 70 K Stage

## Annex B: Cryogenics and Optics Meeting Agenda

### ALMA Front-end Cryogenics and Optics Meeting

Rutherford Appleton Laboratory

Conference Room 12, Building R68

Wednesday 26<sup>th</sup> September, 2001

9:00	Welcome	RAL
9:05	Aims and objectives of front-end meeting	Wolfgang Wild
9:15	Front-end noise gain budget	Gie Han Tan

#### Cryogenics Meeting:

9:25	Review of cryostat requirements and specifications	Anna Orłowska
9:35	Cryostat mechanical and thermal design, including report on thermal link performance Receiver cartridge designs inc. allowable component mass & power budget Prototype construction and test schedule	Anna Orłowska /Mark Harman
11:00	Preliminary open discussion on cryostat design	All

COFFEE BREAK – PLUS ...

10.45

#### Receiver Optics Meeting:

11:00	Overview of current optics design & specifications	Matt Carter
11:10	Cartridge designs <i>These presentations on the different bands must be limited to their impact on the cryostat. The times indicated are maximum times. These should include any changes in the optics from the Lamb report, weight, tolerances LO, etc.. It would be appreciated if as much information as possible can be circulated before the meeting.</i> Bands 1 and 2 and common optics Band 3, 6 and OMTs Bands 4, 8 and 10 Band 7 Band 9 Local oscillator and amplifier requirements	Matt Carter John Payne Yutaro Sekimoto Stefan Claude Andrey Barychev John Webber
12:15	Preliminary optics discussion period	All

BUFFET LUNCH

13:00		
14:00	ALMA band optics calculations	Choy Y. Tham

14:15	New Tolerance budget	Bernard Lazareff
14:25	WVR	Richard Hills
14:35	Calibration issues	Matt Carter
14:45	Calibration at ASTE	Kotaro Kohno
14:55	Windows and filters. Discussion with contributions from Ediss,Carter,Lazareff, Navarro and Yoong Tham	

### Coffee Break

15:45

16:00	Open discussion to re. cryostat and optic design issues. For example, LO requirements, cryostat windows, interfaces, integration and assembly.	All
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17:15	Performance of cryocoolers	Hideo Ogawa
	Progress on Japanese cryostat	Yutaro Sekimoto

17:45	Summary of the day	Wild
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18:00	FINISH AND TRANSPORTATION TO ACCOMMODATION	
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<b>19:30</b>	<b>Evening reception at the Cosener's House, Abingdon.</b>	
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**ALMA Front-end Cryogenics and Optics Meeting**  
**Rutherford Appleton Laboratory**  
**Conference Room 3, Building R61**

**Thursday 27<sup>th</sup> September**

9:00	Review of previous day's contributions and discussions	Matt Carter Anna Orłowska
9:45	Cryostat: Resolution of outstanding issues and concerns relevant to and affecting construction of the pre-production prototype.	Anna Orłowska Matt Carter
	<b>COFFEE BREAK</b>	
10:30		
10:45	continuation	
11:45	Future work for the optics and cryogenics groups	Anna Orłowska Matt Carter
12:15	Resume of meeting	Wolfgang Wild
	<b>LUNCH</b>	
12:30		
13:30	<b>LAB VISITS</b>	
14:00	<b>MIXER MEETING</b>	Bernard Lazareff
18:00	<b>FINISH AND TRANSPORTATION TO ACCOMMODATION</b>	

The above session will take place in parallel with the Photonics Workshop which will be located in CR12, R68. Participants of the Cryogenics and Optics meetings are welcome to attend the Photonics Workshop.

## Annex C: Cryogenics and Optics Meeting Attendees and e-mail addresses

### Attendees (some of them only the first day):

Alessandro Navarrini <navarrin@iram.fr>,  
Andrey Baryshev <A.M.Baryshev@sron.rug.nl>,  
Anna Orłowska <A.H.Orłowska@rl.ac.uk>,  
Bernard Lazareff <lazareff@iram.fr>,  
Brian Ellison <b.ellison@rl.ac.uk>,  
Charles Cunningham <Charles.Cunningham@nrc.ca>,  
Choy Yoong Tham <cytham@mrao.cam.ac.uk>,  
Geoff Ediss <gediss@nrao.edu>,  
Gie Han Tan <ghtan@eso.org>,  
Hideo Ogawa <Ogawa@rishiri.cias.osakafu-u.ac.jp>,  
Inatani Junji <inatani.junji@nasda.go.jp>,  
Jacob Baars <jbaars@eso.org>,  
John Payne <jpayne@nrao.edu>,  
Kotaro Kohno <kkohno@ioa.s.u-tokyo.ac.jp>,  
Marc Rafal <mrafal@nrao.edu>,  
Mark Harman <MRH67@rl.ac.uk>,  
Masato Ishiguro <ishiguro@nro.nao.ac.jp>,  
Matt Carter <carter@iram.fr>,  
Noguchi Takashi <noguchi@nro.nao.ac.jp>,  
Renzo Nesti <nesti@arcetri.astro.it>  
Richard Hills <richard@mrao.cam.ac.uk>,  
Richard Kurz <rkurz@eso.org>,  
Richard Simon <rsimon@nrao.edu>,  
Robert Brown <rbrown@nrao.edu>,  
Ronald Hesper <hesper@sron.rug.nl>,  
Stephane Claude <claud@iram.fr>,  
Tetsuo Hasegawa <tetsuo.hasegawa@nao.ac.jp>  
Ueda Akitoshi <a.ueda@nao.ac.jp>,  
Victor Belitsky <belitsky@oso.chalmers.se>,  
Wolfgang Wild <wild@astro.rug.nl>,  
Yutaro Sekimoto <sekimoto@nro.nao.ac.jp>

### Unable to attend:

Eric Bryerton <ebryerto@nrao.edu>,  
Gene Lauria <glauria@nrao.edu>,  
James Lamb <lamb@ovro.caltech.edu>,  
John Webber <jwebber@nrao.edu>,  
Karl Schuster <schuster@iram.fr>,  
Richard Plambeck <plambeck@astron.berkeley.edu>,  
Ryohei Kawabe <kawabe@nro.nao.ac.jp>,  
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