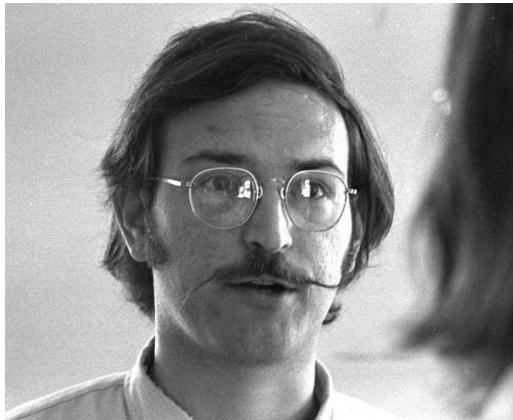


# 4He Abundances: Optical vs Radio

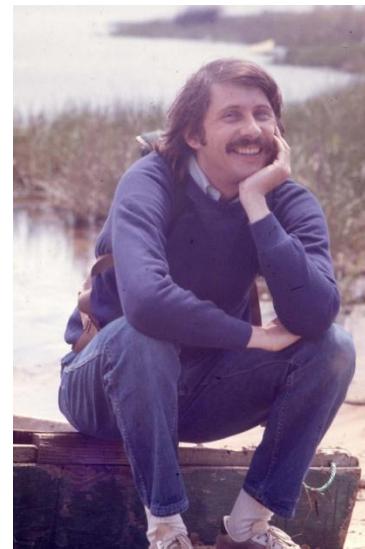
Dana S. Balser



# Collaborators



Tom Bania

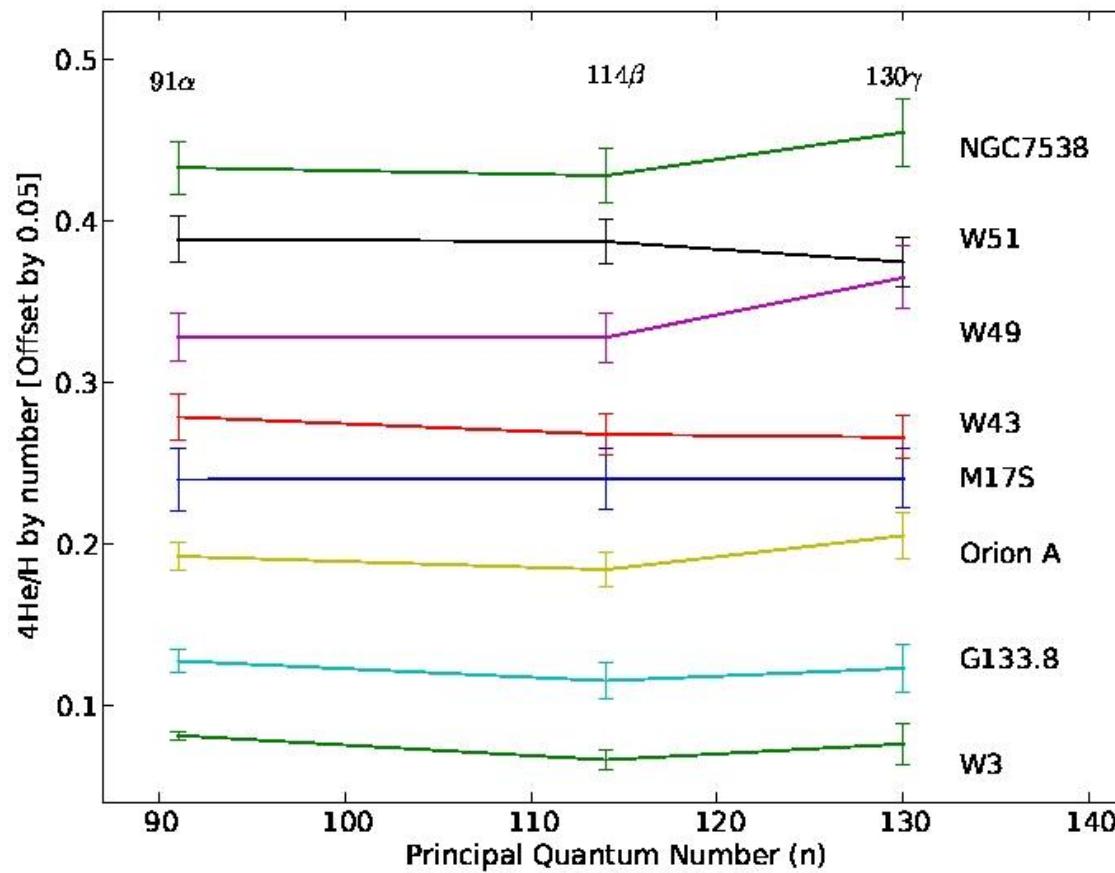


Bob Rood

## Radio Recombination Lines

- Sensitivity
- Spectral Baselines
- Departures from LTE
- Ionization Structure
- Density Structure

# ${}^4\text{He}/\text{H}$ : RRLs with the 140 Foot Telescope



Peimbert et al. (1992)

# Green Bank Telescope Observations



Planetary Nebulae :

NGC 3242

NGC 6543

NGC 6826

NGC 7009

HII Regions :

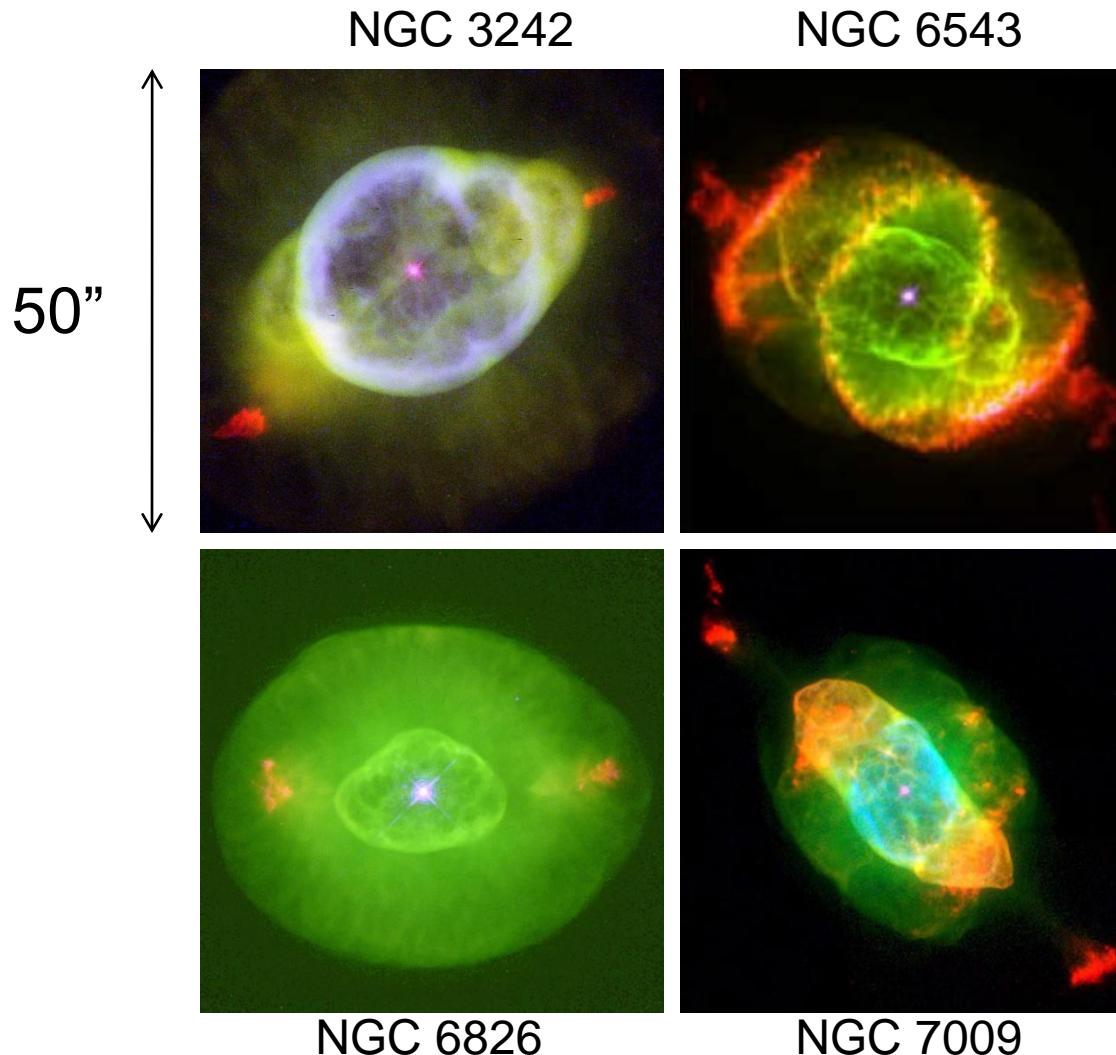
M17

S206

GBT HPBW = 80" at 9 GHz

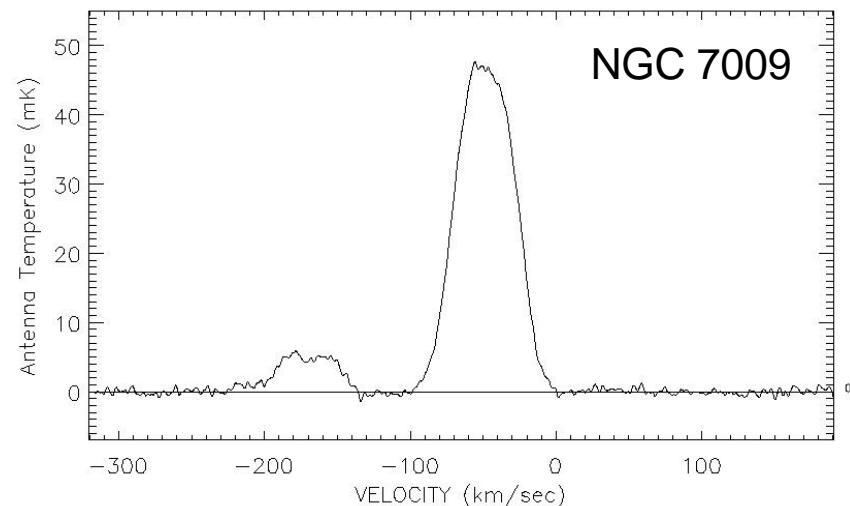
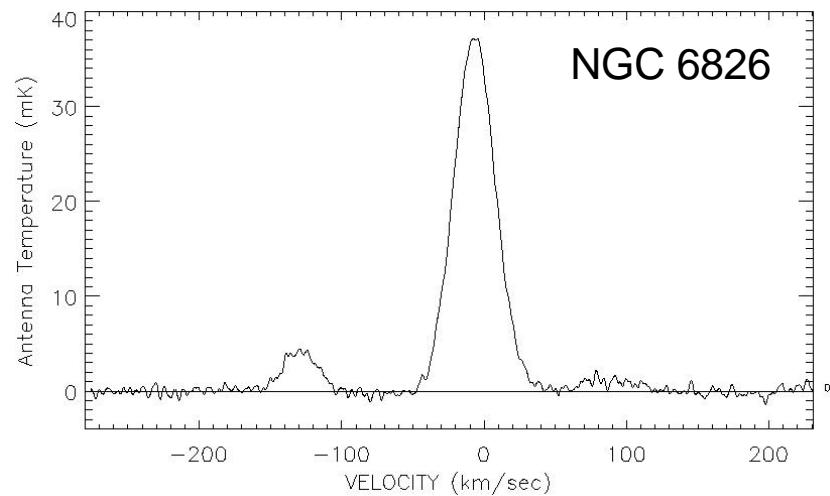
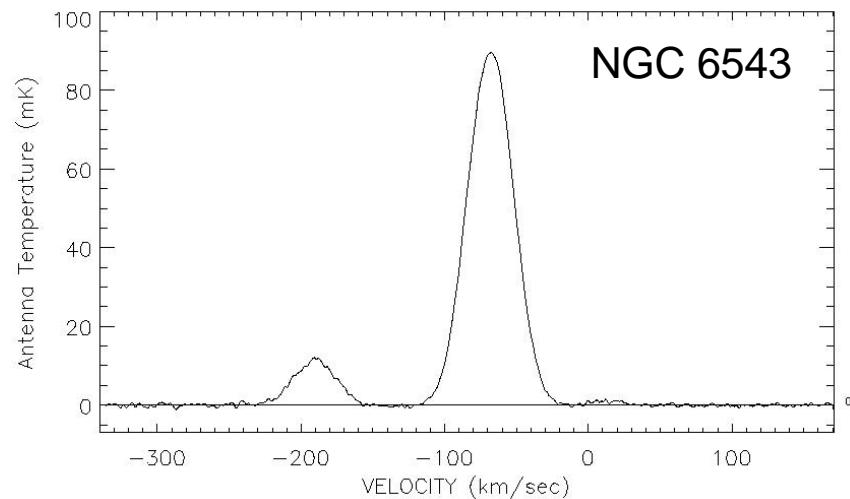
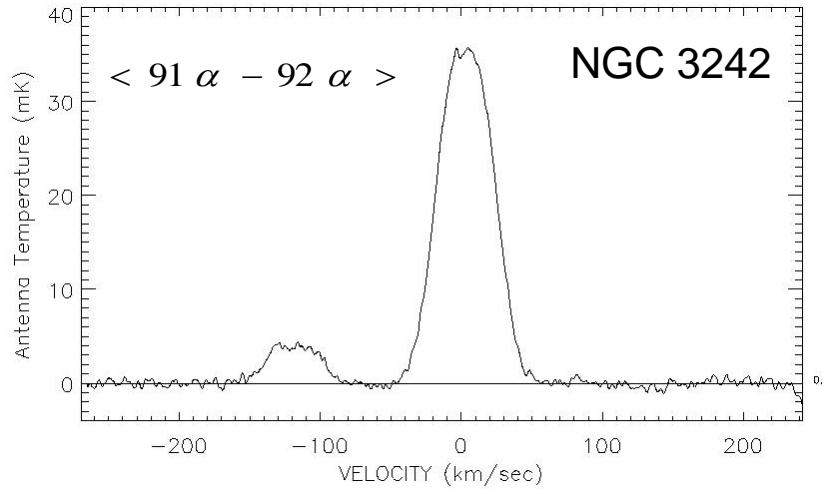
H and He 87  $\alpha$  – 93  $\alpha$

# Planetary Nebulae



Balick et al.

# GBT: PNe Radio Recombination Lines



# HII Regions: M17

Optical (Halpha)

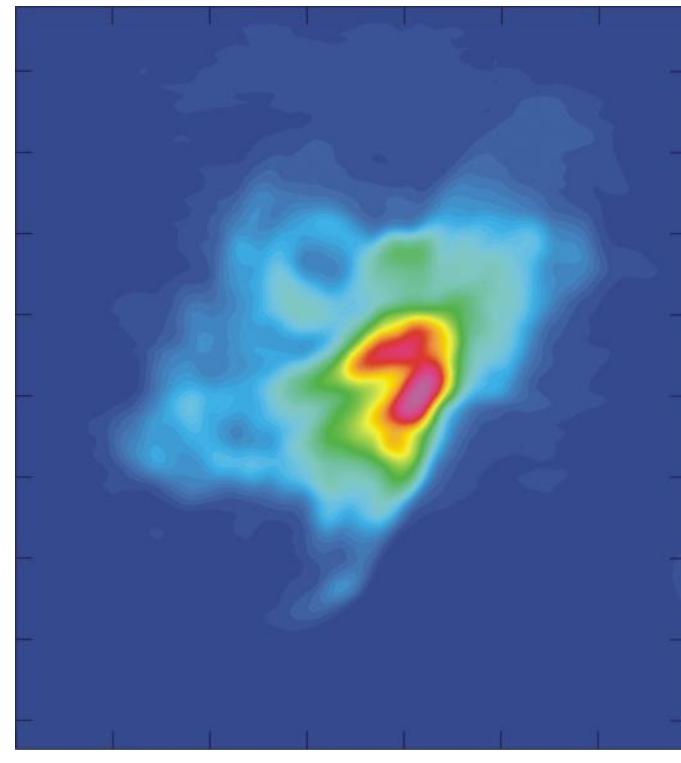


↔

30'

MacQuarrie

Radio Continuum (9GHz)



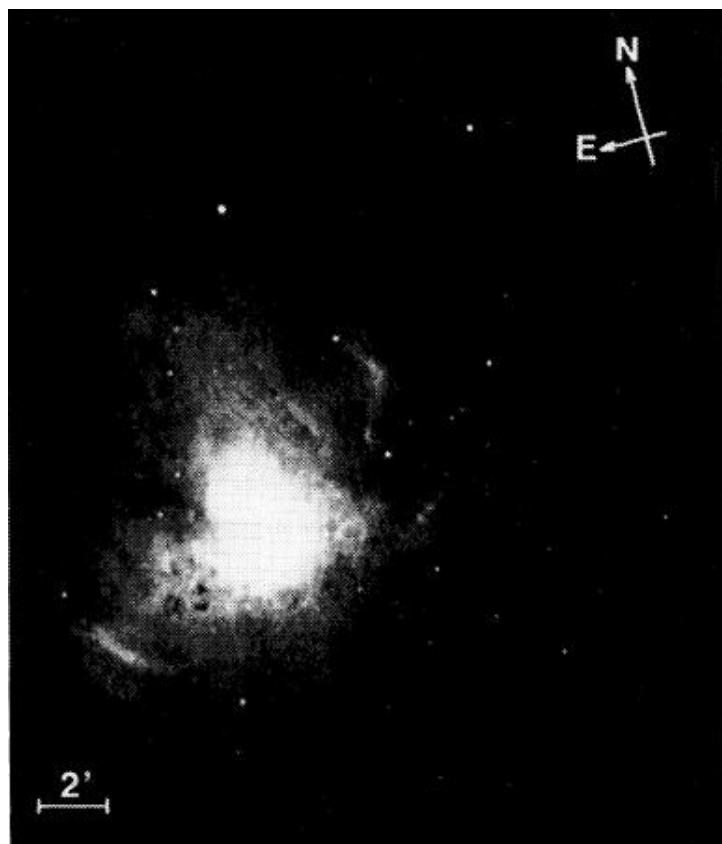
↔

60'

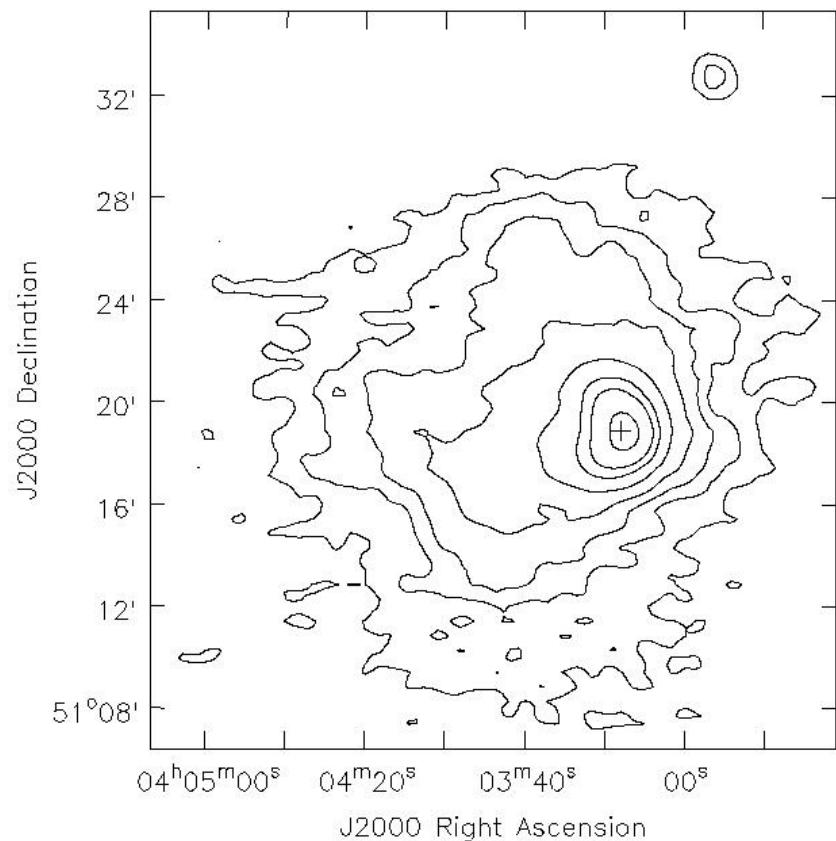
Maddalena et al.

# HII Regions: S206

Optical (Halpha)



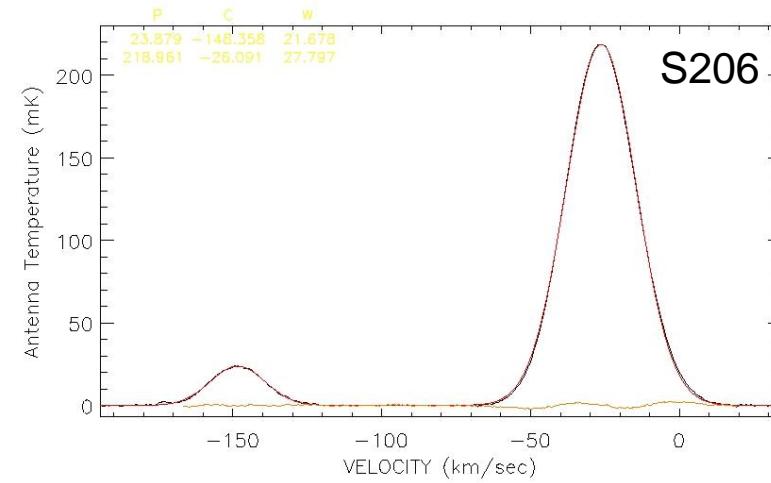
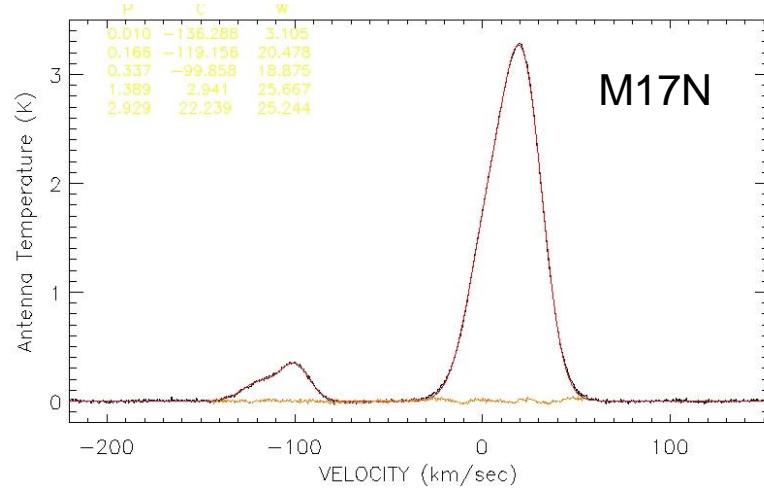
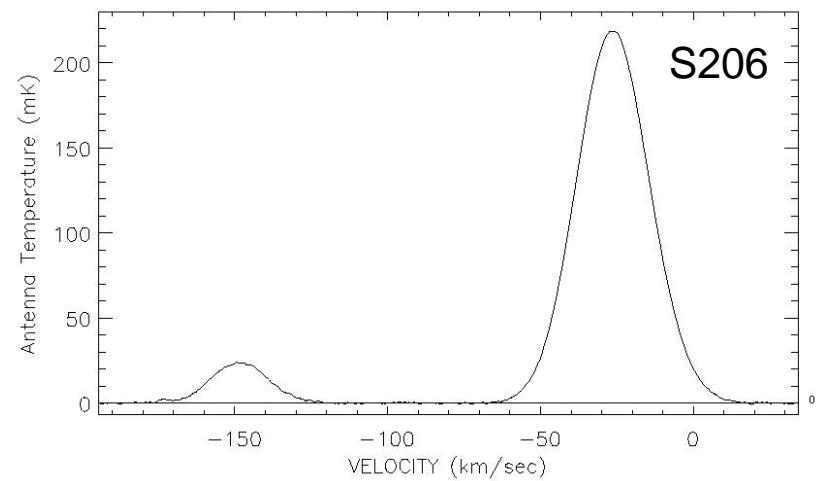
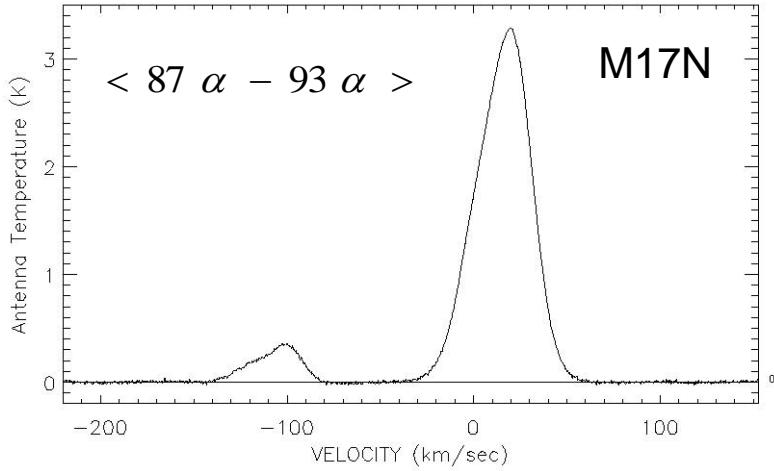
Radio Continuum (9 GHz)



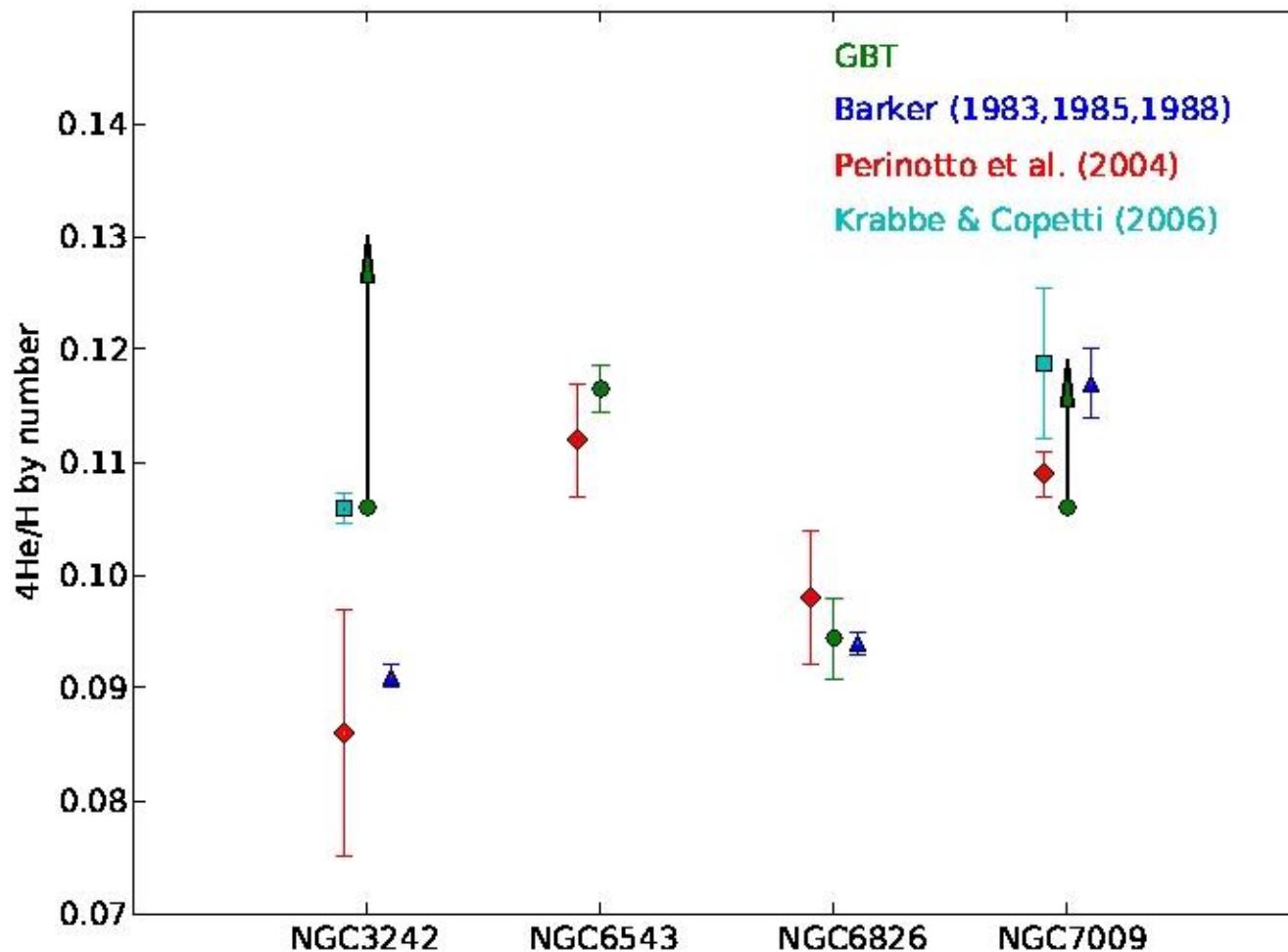
Pismis & Mampaso (1991)

Balser (2006)

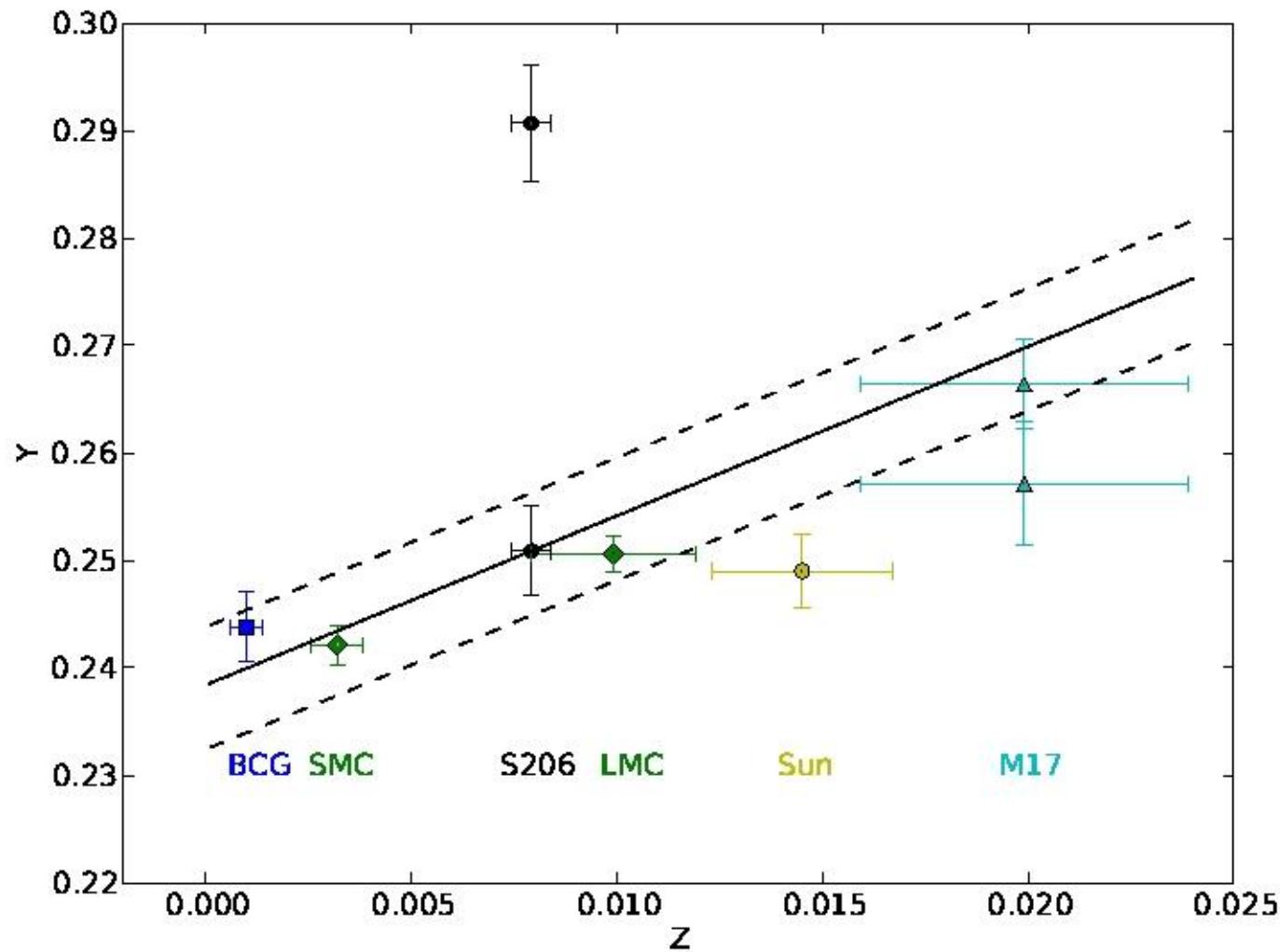
# GBT: HII Region Radio Recombination Lines



# ${}^4\text{He}/\text{H}$ : Planetary Nebulae



# $^{4}\text{He}/\text{H}$ : HII Regions



## Summary

- ${}^4\text{He}/\text{H}$  differences as large as 20% between Optical and Radio
- Galactic  $dY/dZ \sim 1$  (M17, S206)

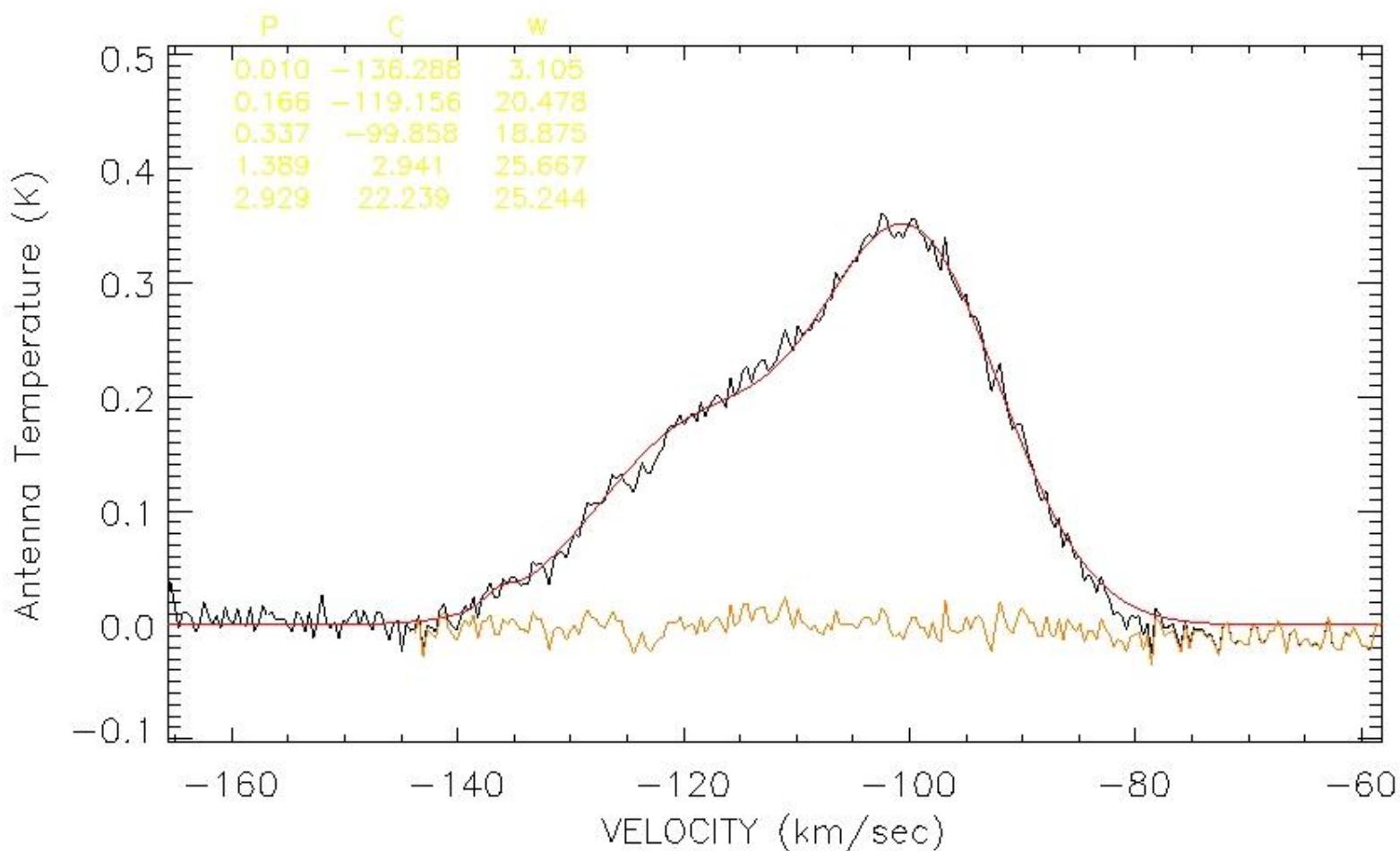
## Future Work

- Explore  ${}^4\text{He}/\text{H}$  with  $n$
- Models of PNe and HII Regions
- EVLA

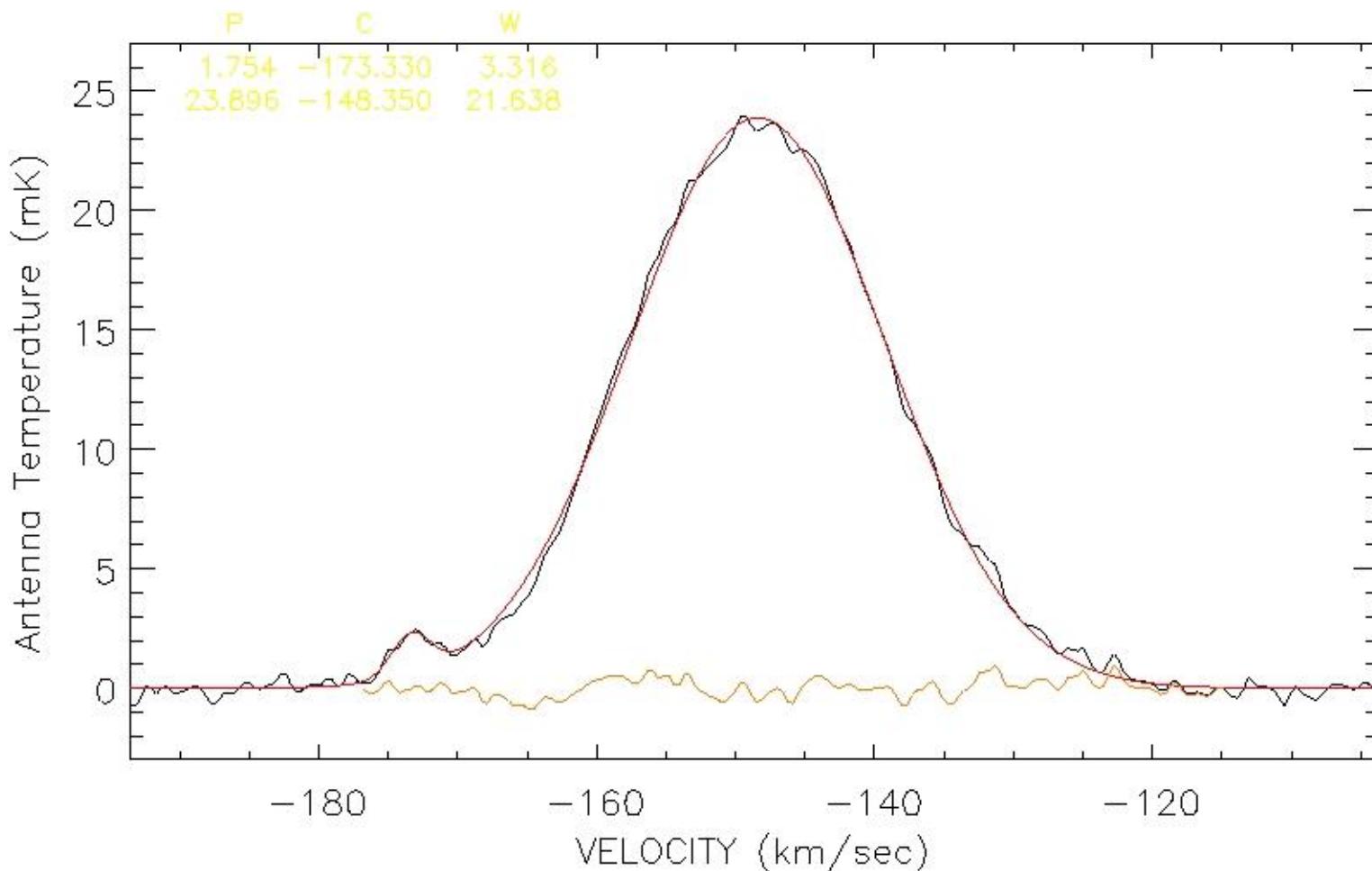


# Extra Slides

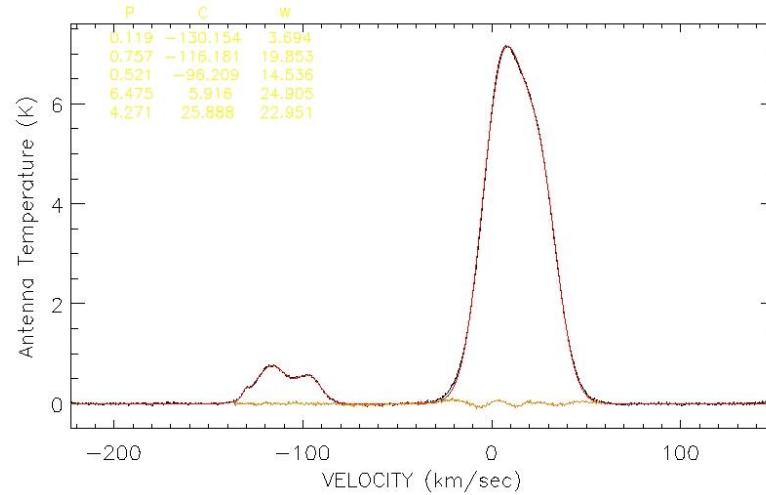
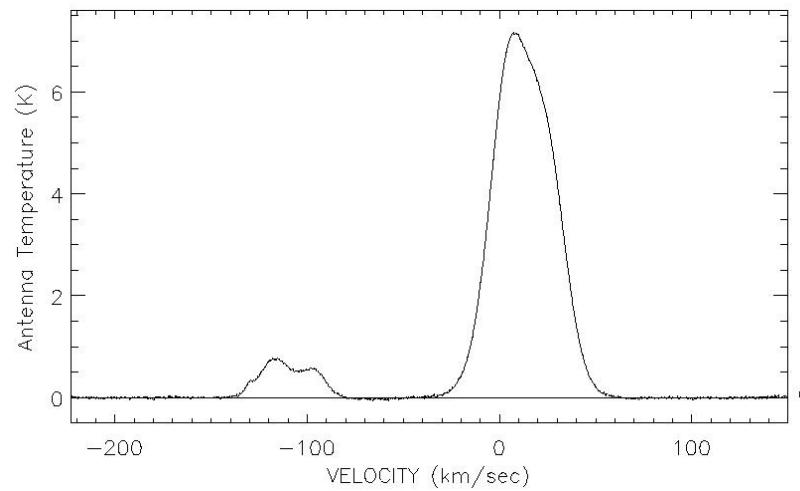
# Expanded View: M17N



## Expanded View: S206



# M17S



## Expanded View: M17S

