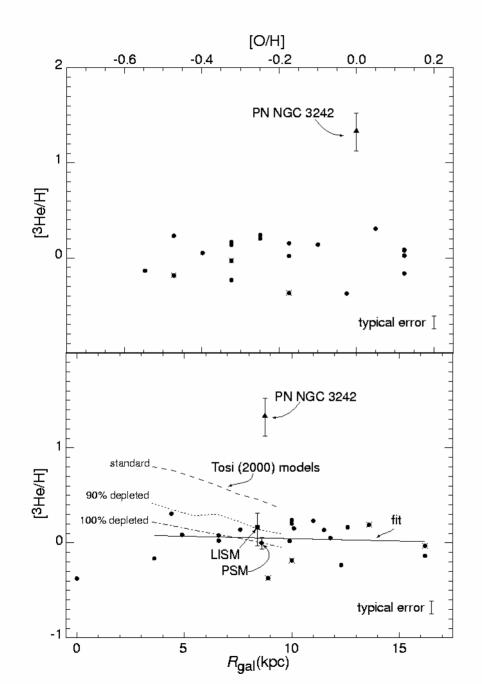
# Helium-3 in Planetary Nebulae

Rood, Bania (BU), Balser (NRAO), Goss (NRAO), Quireza (ON, Brazil), Wilson (MPIfR) Observe 3He using the hyperfine (spin-flip) line of 3He+

Analog of the 21 cm line of H

$$v = 8665.65 \text{ MHz}$$

$$\lambda = 3.36$$
 cm



#### Bania, Rood, & Balser 2002

$$\eta_{10} = 5.4^{+2.2}_{-1.2}$$
 
$$\Omega_B = 0.04$$

### Spergel et al. 2003, WMAP

$$\begin{split} \eta_{10} &= 6.5^{+0.4}_{-0.3} \\ \Omega_B &= 0.047 \pm 0.006 \end{split}$$

One is not enough!

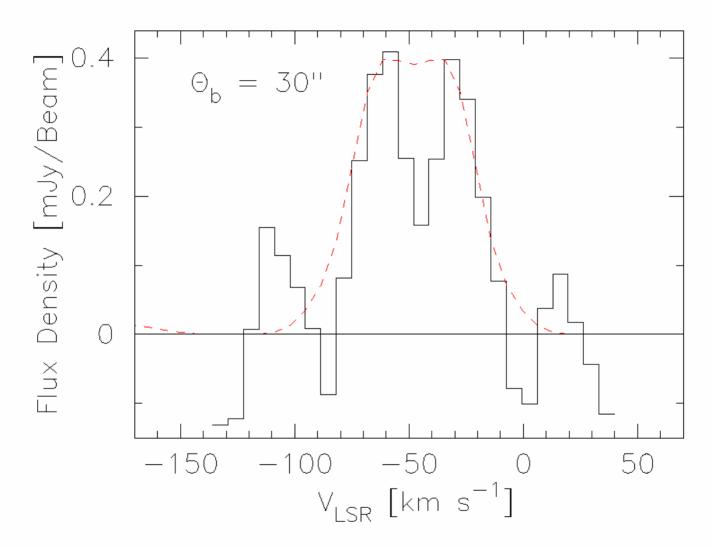
Except in cosmology

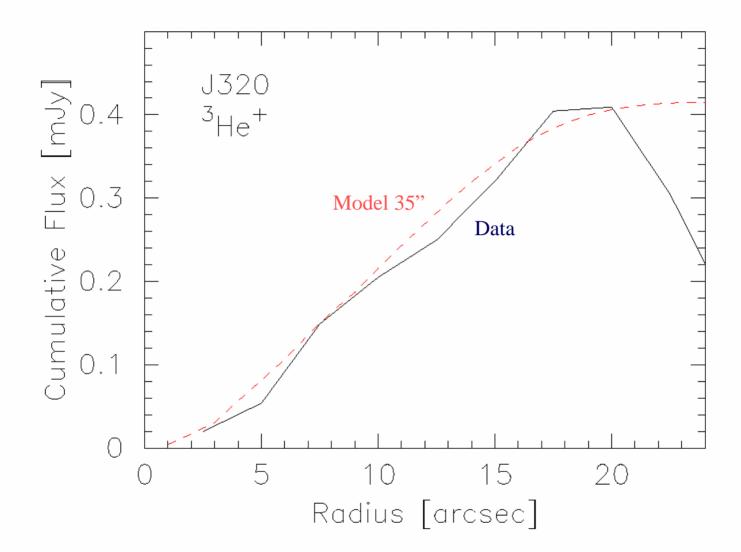
The PN sample:

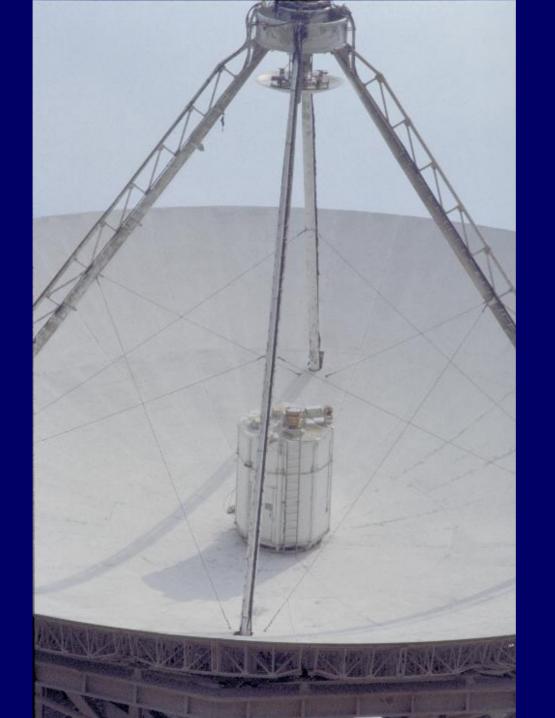
Why should I read a slide to you?



PNe He3 at the VLA: Balser, Goss, Bania, Rood (2005)





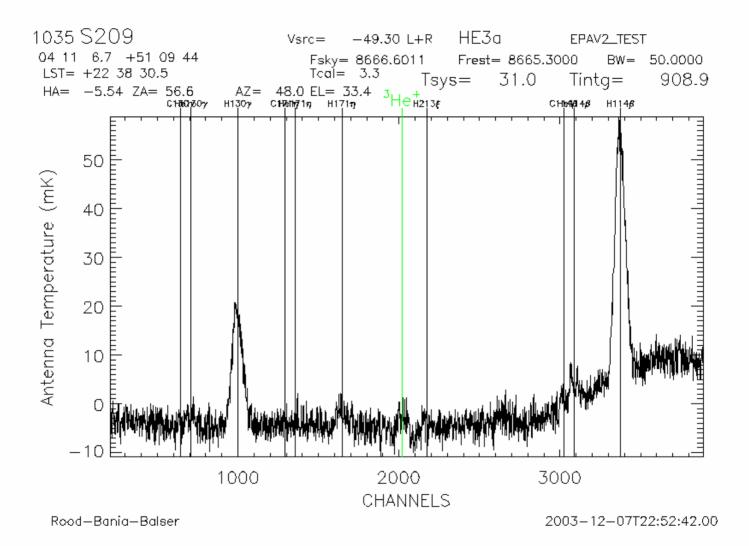


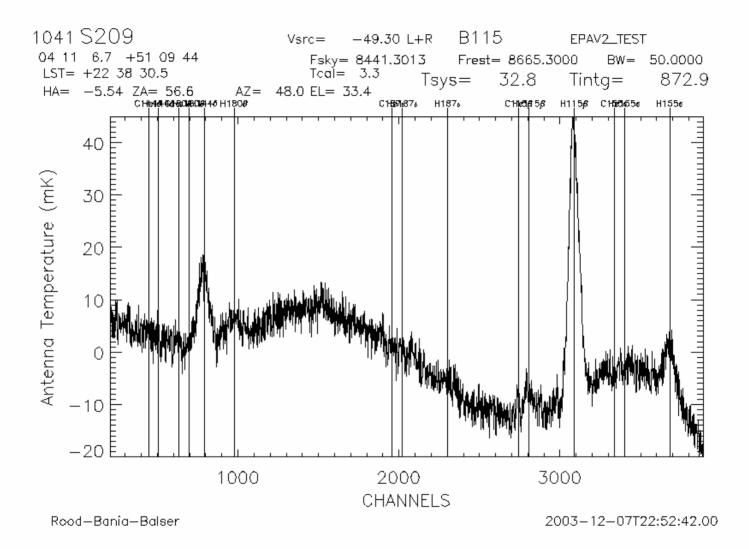


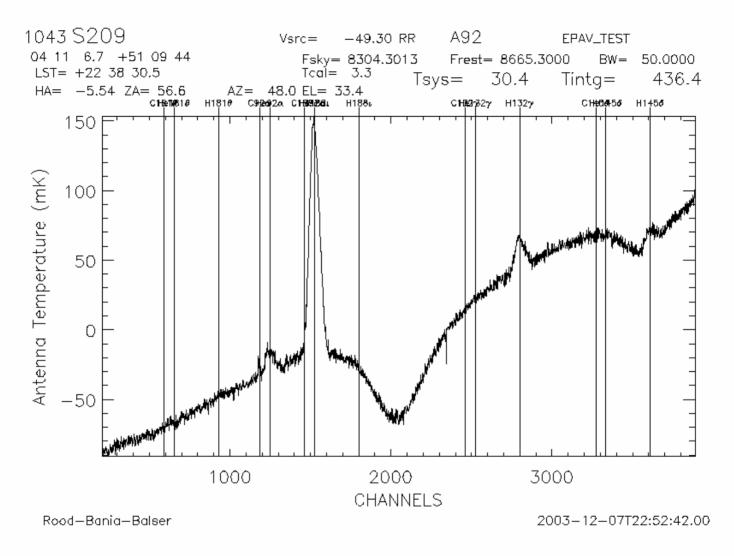




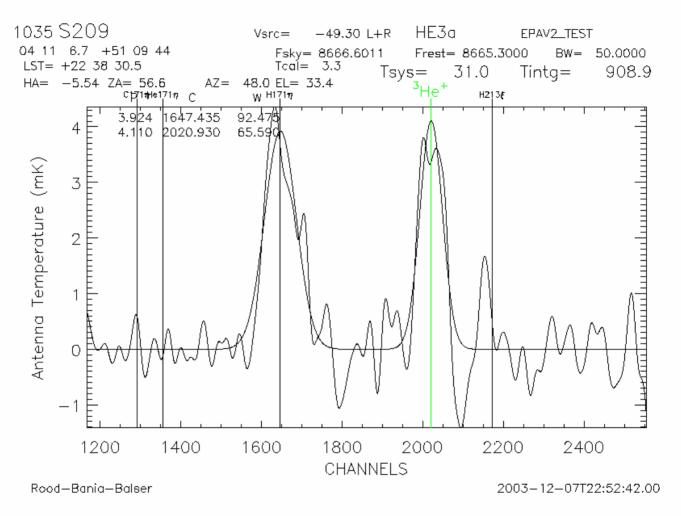




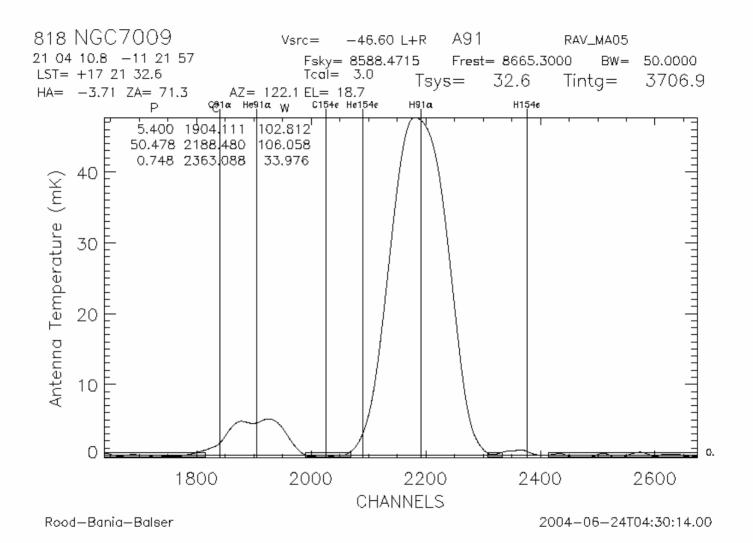


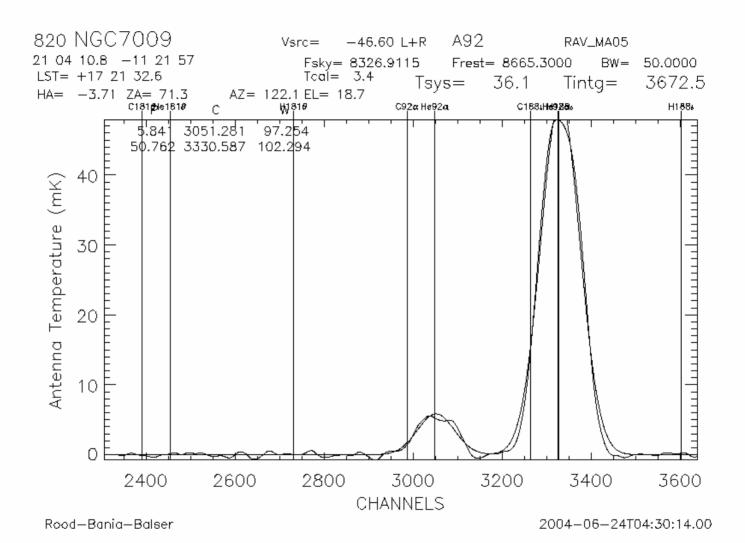


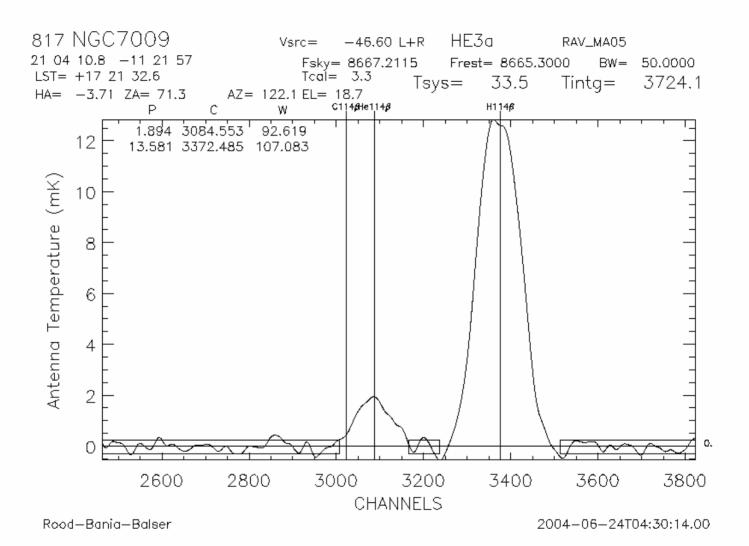
Some days it's chicken; some days it's feathers

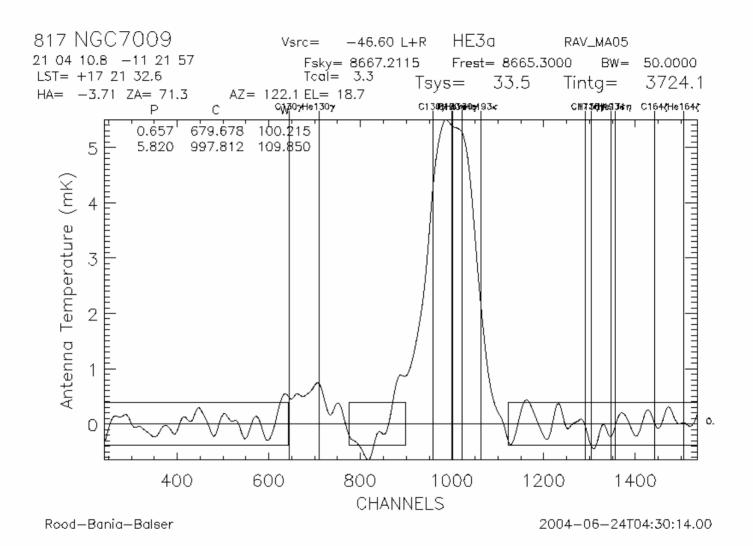


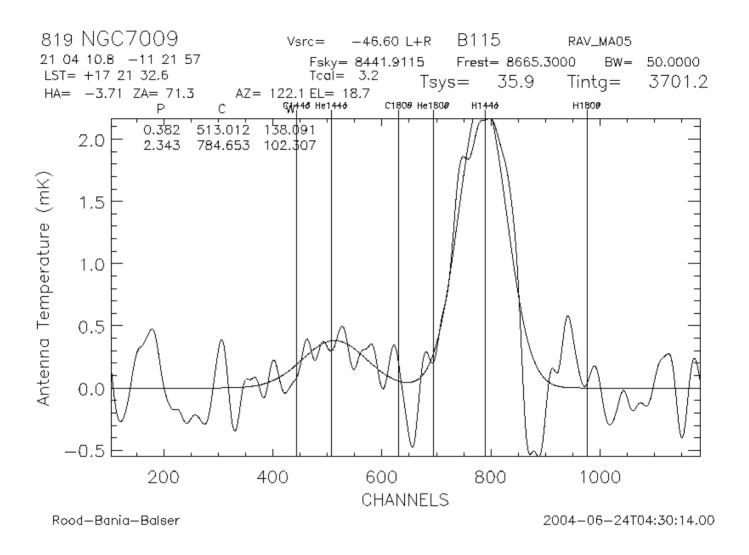
He3 in S209 in only 7.5hr!

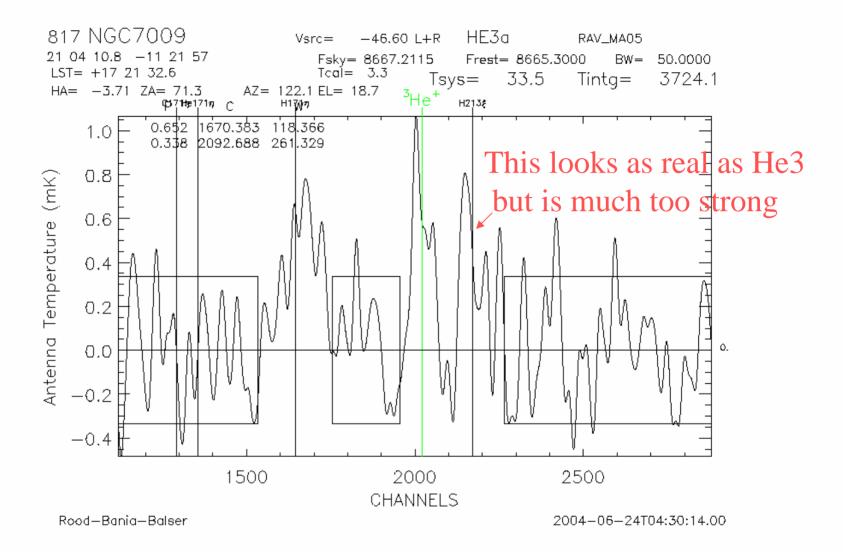


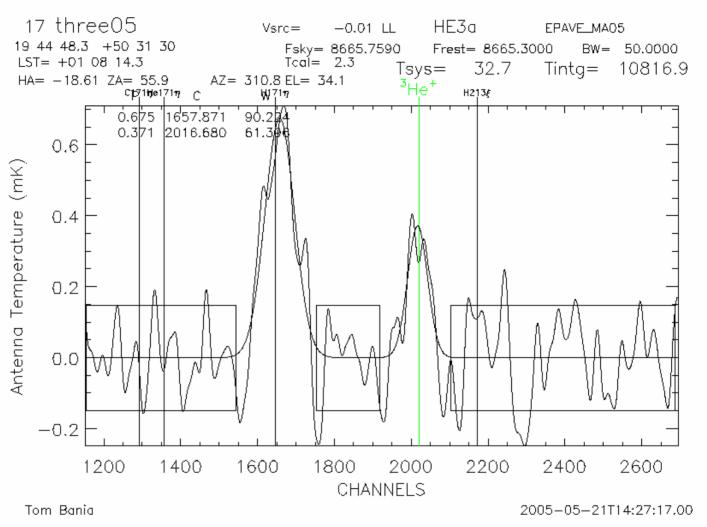




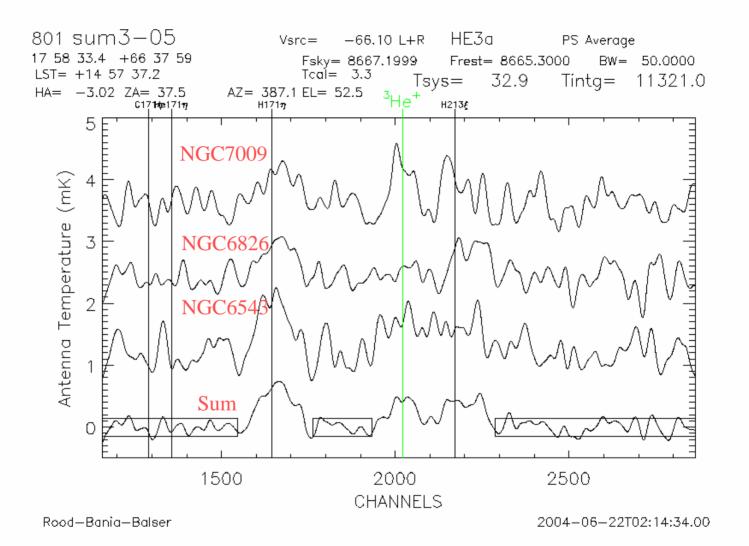








NGC7009 + NGC6543 + NGC6826



### **GBT Conclusions**

- Standing waves are not a problem
- There is still baseline structure (BS) probably resulting from mismatches in the IF system.
  - BS varies with frequency sometimes almost invisible other times very problematic
  - BS amplitude is proportional to source continuum and moves with sky frequency
- At the mK level there are pseudo-lines
- In some AC bands there are short duration spikes in the ACF at seemingly random times, lags, and amplitudes

## Helium-3 Conclusions

- We have found helium-3 in another PN, J320, using the VLA
- We probably have found helium-3 in NGC7009 using the GBT and may have a second detection in NGC6543
- The scheduling mode and proposal pressure on the GBT may not allow us to solidify these results in the near future.

#### A bonus: He<sup>++</sup> or O<sup>++</sup> RRL (a first?)

